

Scalar Dark Matter and the Higgs Portal



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Theoretical Nuclear, Particle, Astrophysics & Cosmology

<http://www.physics.wisc.edu/groups/particle-theory/>

Snowmass EW Workshop,
Duke, February 2013

Goals

- *Motivate Higgs portal DM*
- *Introduce generic models*
- *Outline signatures & constraints:
modified Higgs properties, new states,
EWPO, vacuum stability*
- *Discuss future directions*

Outline

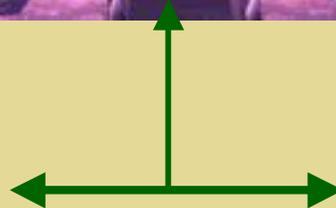
1. *Dark Matter Portals*
2. *Why the Higgs portal?*
3. *Why scalar dark matter ?*
4. *General considerations*
5. *Simplest examples: scalar DM & LHC*
6. *Discussion*

I. Dark Matter Portals



Standard Model

Hidden Sector (DM)



Portals

Two approaches to particle dark matter :

- *Specific model (MSSM....)*
- *“Model independent”*

Model Independent Portals

- *Vector portal (“dark photons”...)*
- *Neutrino portal*
- *Axion portal*
- *Higgs portal*
- *Higher dimensional op’s portal*

Model Independent Portals

- *Vector portal* (“dark photons”...)
- *Neutrino portal*
- *Axion portal*
- *Higgs portal*
- *Higher dimensional op’s portal*

$$\mathcal{L}_4 = \Lambda^4 + \mu^2 \mathcal{O}_2 + \mathcal{O}_4 + \frac{1}{\Lambda_{\text{BSM}}} \mathcal{O}_5 + \dots$$

Model Independent Portals

- *Vector portal* (“dark photons”...)
- *Neutrino portal*
- *Axion portal*
- *Higgs portal*
- *Higher dimensional op’s portal*

$$\mathcal{L}_4 = \Lambda^4 + \mu^2 \mathcal{O}_2 + \mathcal{O}_4 + \frac{1}{\Lambda_{\text{BSM}}} \mathcal{O}_5 + \dots$$

Higgs Portal

$$\mathcal{O}_4 = \lambda_{\phi H} \phi^\dagger \phi H^\dagger H$$

- *Renormalizable*
- *Z_2 symmetric*
- *Dimensionless coupling*
- *ϕ (DM): singlet or charged
under $SU(2)_L \times U(1)_Y$*

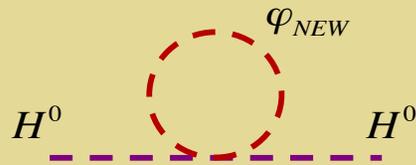
II. Why the Higgs Portal ?



Scalar Fields in Particle Physics

Scalar fields are a simple

Scalar fields are theoretically problematic



$$\Delta m^2 \sim \lambda \Lambda^2$$

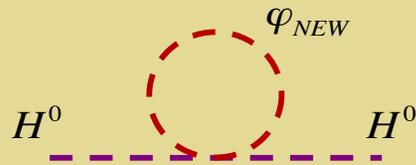
Has a fundamental scalar finally been discovered ?

If so, is it telling us anything about Λ ?

Scalar Fields in Particle Physics

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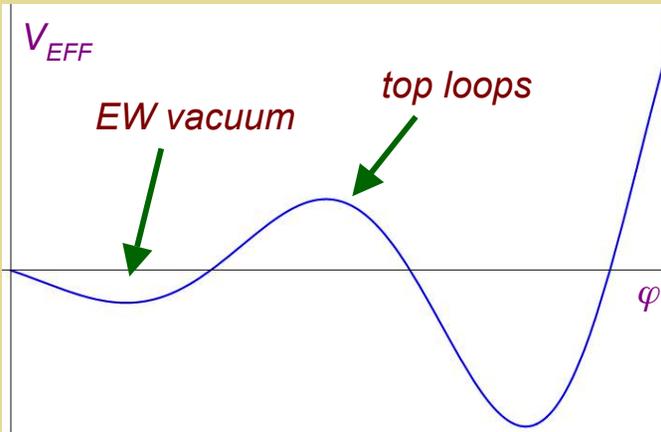
$$\Delta m^2 \sim \lambda \Lambda^2$$

Must $\Lambda_{BSM} \sim \text{TeV}$ to maintain a weak scale scalar ?

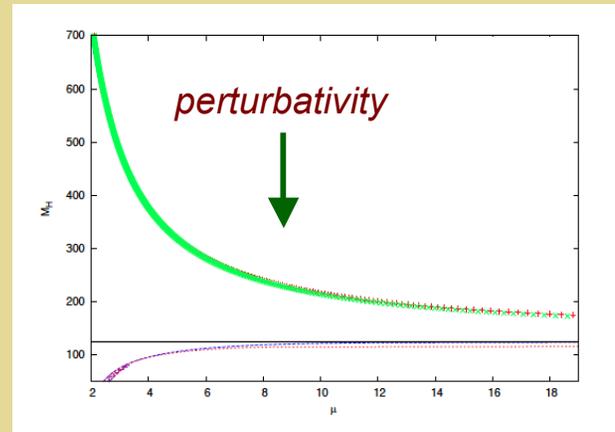
Perhaps new weak scale physics couples only to scalar sector: “Higgs portal”

Stable EW Vacuum ?

Preserving EW Min



"Funnel plot"



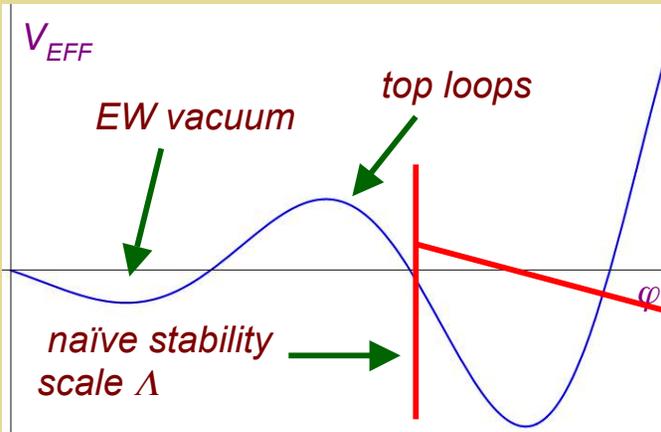
$$\beta_\lambda = \frac{1}{16\pi^2} \left(4\lambda^2 - 36y_t^4 + 12\lambda y_t^2 - 9\lambda g^2 - 3\lambda g'^2 + \frac{9}{4}g'^4 + \frac{9}{2}g^2 g'^2 + \frac{27}{4}g^4 \right)$$

sets m_H

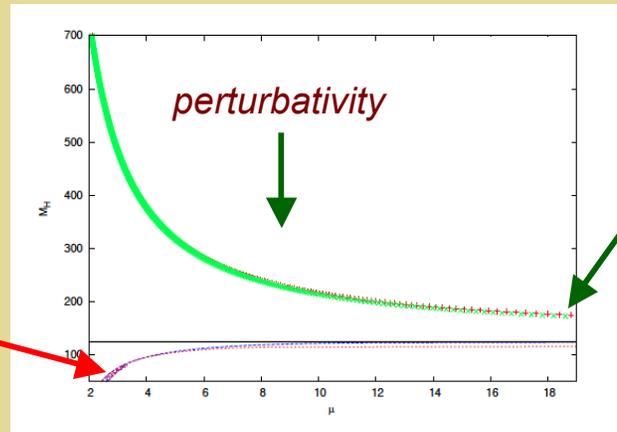
top loops

Stable EW Vacuum ?

Preserving EW Min



“Funnel plot”



SM stability
& pert'vity

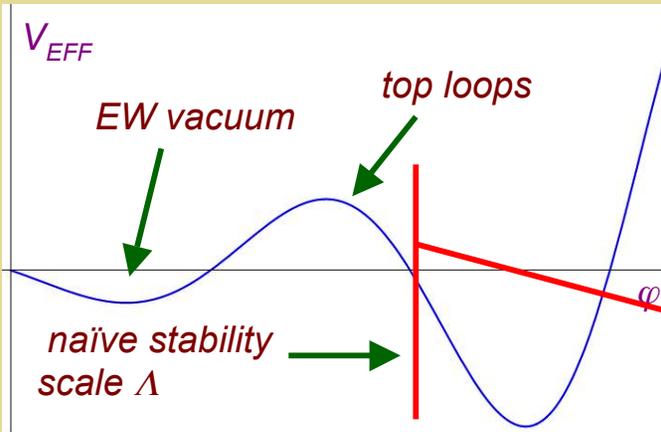
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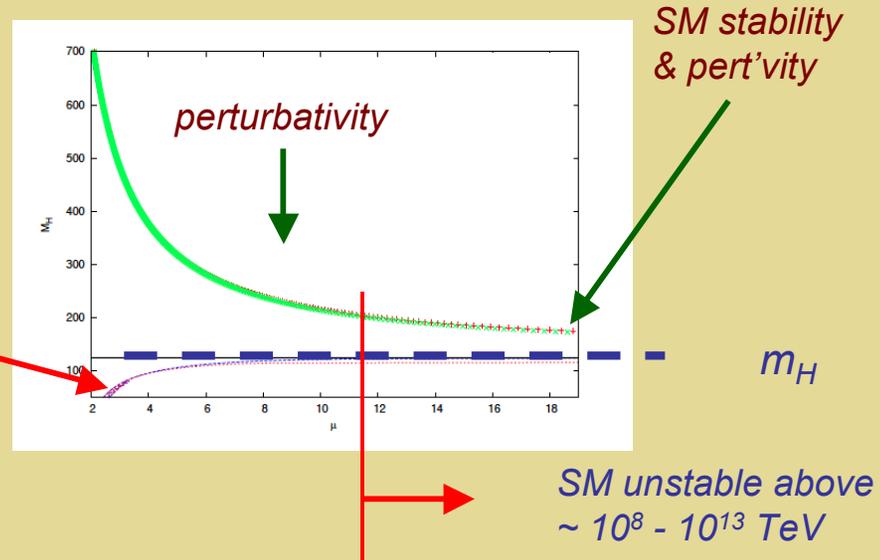
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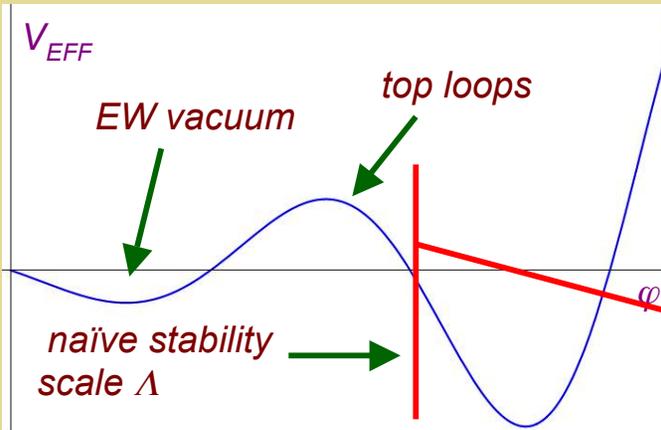
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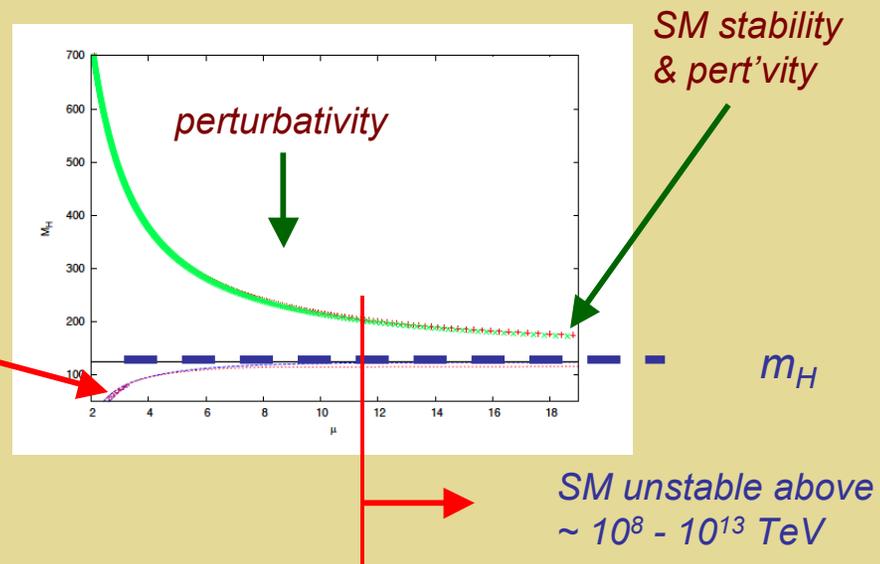
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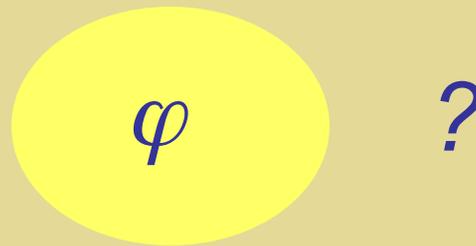
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sets m_H

top loops

Higgs portal interactions \rightarrow more robust stability

III. Why scalar dark matter ?



Scalar Fields in Cosmology

*What role do scalar fields play (if any)
in the physics of the early universe ?*

Scalar Fields in Cosmology

<i>Problem</i>	<i>Theory</i>	<i>Exp't</i>
<ul style="list-style-type: none">• <i>Inflation</i>• <i>Dark Energy</i>• <i>Dark Matter</i>• <i>Phase transitions</i>		

Scalar Fields in Cosmology

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• <i>Phase transitions</i>	✓	?

- *Could experimental discovery of a fundamental scalar point to early universe scalar field dynamics?*
- *Are there signatures in modified Higgs properties, new states, or EW precision tests ?*

Scalar Fields in Cosmology

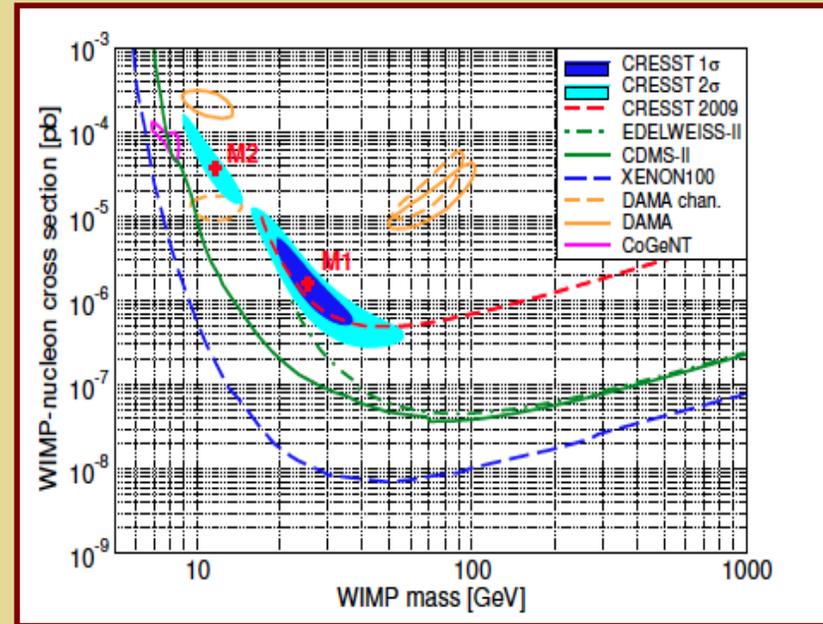
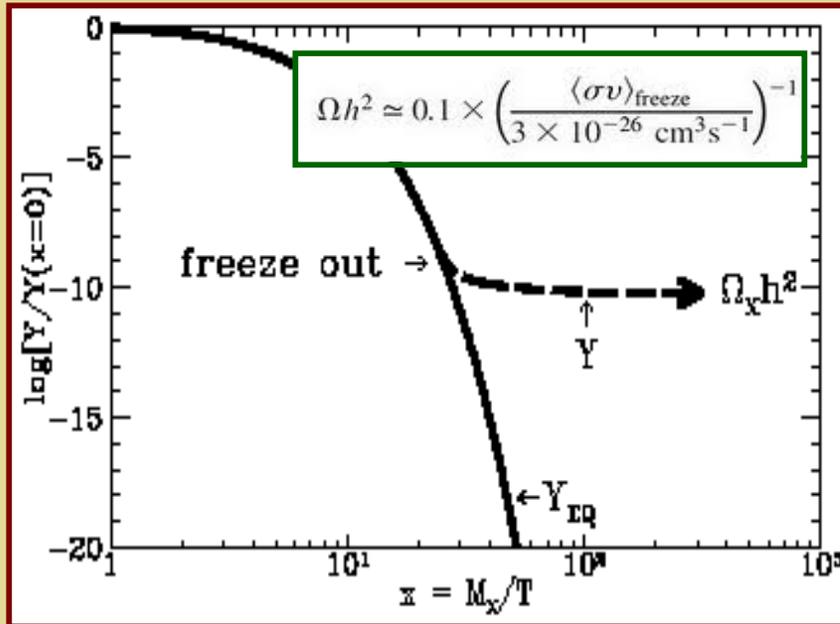
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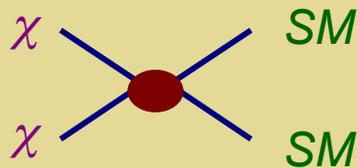
Focus of this talk, but perhaps part of larger role of scalar fields in early universe

IV. General Considerations

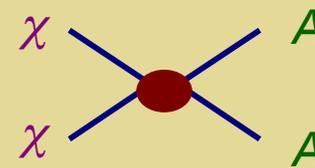
Thermal DM: Ω_{CDM} & σ_{SI}



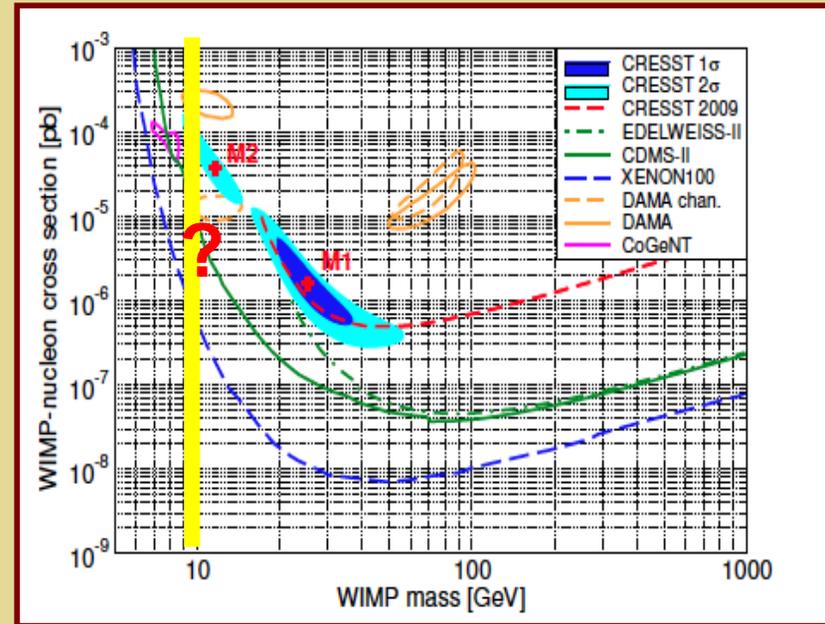
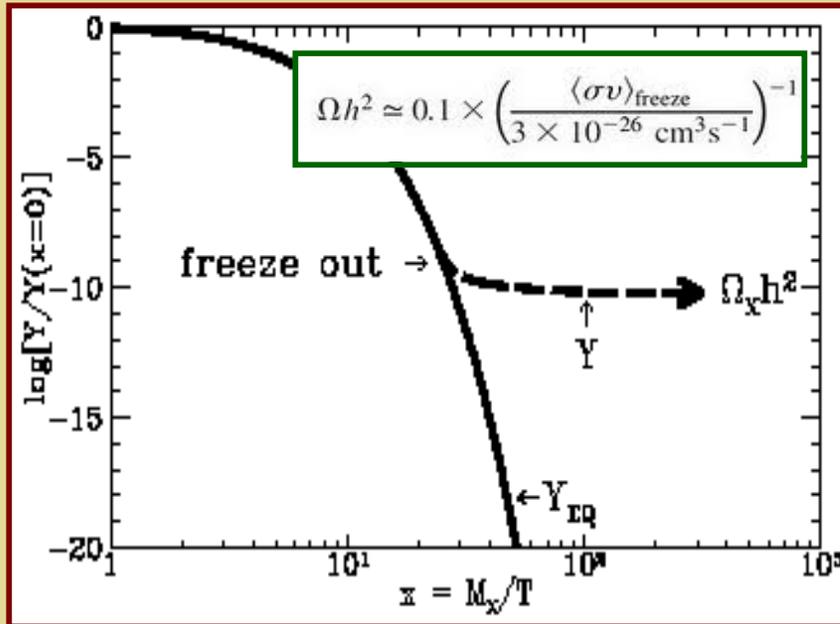
Thermal DM: WIMP



Direct detection: Spin-indep DM-nucleus scattering

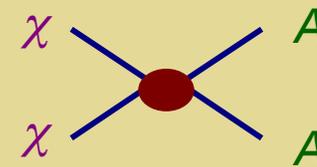
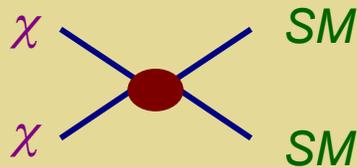


Thermal DM: Ω_{CDM} & σ_{SI}

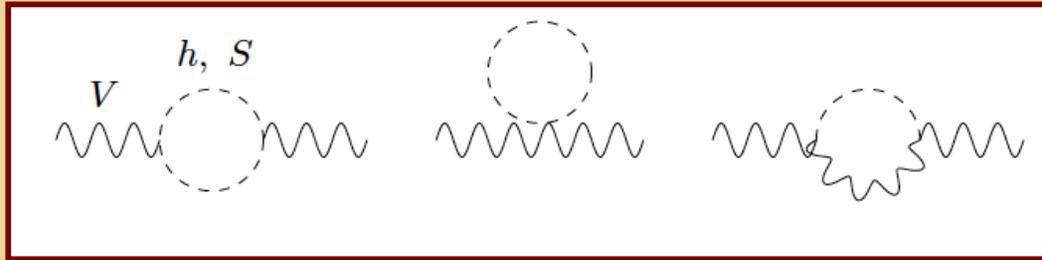


Thermal DM: WIMP

Direct detection: Spin-indep DM-nucleus scattering



EWPO



$$\hat{\alpha}S = \frac{4\hat{s}^2\hat{c}^2}{M_Z^2} \text{Re} \left\{ \hat{\Pi}_{ZZ}(0) - \hat{\Pi}_{ZZ}(M_Z^2) + \frac{\hat{c}^2 - \hat{s}^2}{\hat{c}\hat{s}} \left[\hat{\Pi}_{Z\gamma}(M_Z^2) - \hat{\Pi}_{Z\gamma}(0) \right] + \hat{\Pi}_{\gamma\gamma}(M_Z^2) \right\}^{\text{New}},$$

$$\hat{\alpha}T = \frac{1}{M_W^2} \left\{ \hat{c}^2 \left(\hat{\Pi}_{ZZ}(0) + \frac{2\hat{s}}{\hat{c}} \hat{\Pi}_{Z\gamma}(0) \right) - \hat{\Pi}_{WW}(0) \right\}^{\text{New}},$$

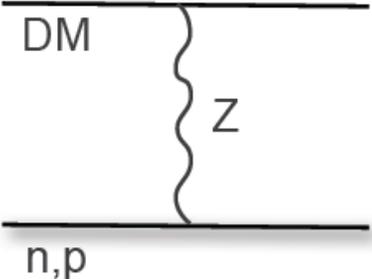
- *New states: h-S mixing*
- *$SU(2)_L \times U(1)_Y$ multiplets: direct contributions*

$$\alpha S \sim \Pi_{3Y}$$

EWPO & σ_{SI}

- *New states: h-S mixing*
- *Gauge interactions: direct contributions*

$$\alpha S \sim \Pi_{3\gamma}$$


$$\sigma_n \sim \frac{g_1^2}{4\pi^2} \frac{m_{n,p}^2}{m_Z^4} \sim 10^{-39} \text{cm}^2$$

$$g_{xxZ} \propto 2 I_3 - 4 Q \sin^2 \theta_W$$

$$Y = 2 (Q - I_3)$$

→ *Want $Y = 0$*

EW corrections break custodial SU(2): $T \neq 0$ (but small)

V. Simple examples

Higgs Portal: Simple Scalar Extensions

<i>Extension</i>	<i>DOF</i>	<i>EWPT</i>	<i>DM</i>
<i>Real singlet</i>	<i>1</i>	<i>✓</i>	<i>✗</i>
<i>Real singlet</i>	<i>1</i>	<i>✗</i>	<i>✓</i>
<i>Complex Singlet</i>	<i>2</i>	<i>✓</i>	<i>✓</i>
<i>Real Triplet</i>	<i>3</i>	<i>✓</i>	<i>✓</i>

May be low-energy remnants of UV complete theory & illustrative of generic features

Higgs Portal: Simple Scalar Extensions

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May be low-energy remnants of UV complete theory & illustrative of generic features

The Simplest Extension

Simplest extension of the SM scalar sector: add one real scalar S (SM singlet)

$$V_{\text{HS}} = \frac{a_1}{2} (H^\dagger H) S + \frac{a_2}{2} (H^\dagger H) S^2$$

$$\text{EWPT: } a_{1,2} \neq 0 \quad \& \quad \langle S \rangle \neq 0$$

$$\text{DM: } a_1 = 0 \quad \& \quad \langle S \rangle = 0$$

O'Connell, R-M, Wise; Profumo, R-M, Shaugnessy; Barger, Langacker, McCaskey, R-M Shaugnessy; He, Li, Li, Tandean, Tsai; Petraki & Kusenko; Gonderinger, Li, Patel, R-M; Cline, Laporte, Yamashita; Ham, Jeong, Oh; Espinosa, Quiros; Konstandin & Ashoorioon...

The Simplest Extension, Cont'd

Mass matrix

$$M^2 = \begin{pmatrix} \mu_h^2 & \mu_{hs}^2/2 \\ \mu_{hs}^2/2 & \mu_s^2 \end{pmatrix}$$

$$\begin{pmatrix} h_1 \\ h_2 \end{pmatrix} = \begin{pmatrix} \sin\theta & \cos\theta \\ \cos\theta & -\sin\theta \end{pmatrix} \begin{pmatrix} h \\ s \end{pmatrix}$$

$$\mu_h^2 \equiv \frac{\partial^2 V}{\partial h^2} = 2\bar{\lambda}_0 v_0^2$$

$$\mu_s^2 \equiv \frac{\partial^2 V}{\partial s^2} = b_3 x_0 + 2b_4 x_0^2 - \frac{a_1 v_0^2}{4x_0}$$

$$\mu_{hs}^2 \equiv \frac{\partial^2 V}{\partial h \partial s} = (a_1) + 2a_2 x_0 v_0$$

$$x_0 = \langle S \rangle$$

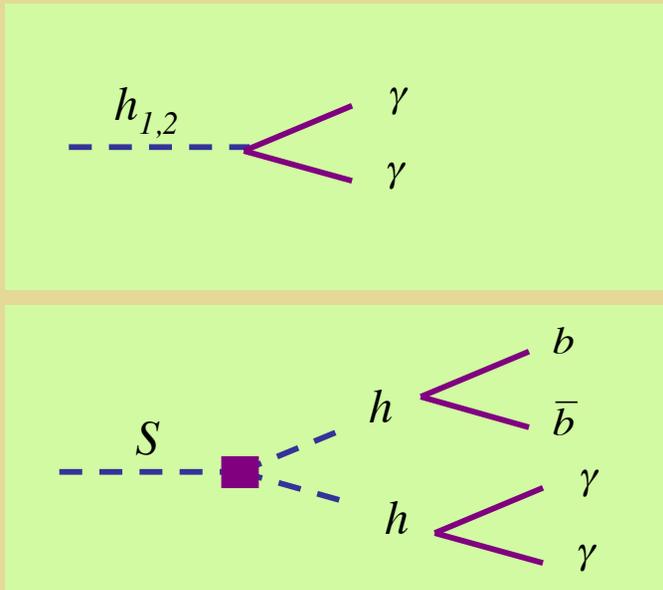
$$\tan\theta = \frac{y}{1 + \sqrt{1 + y^2}},$$

$$y \equiv \frac{\mu_{hs}^2}{\mu_h^2 - \mu_s^2}$$

$$m_{1,2}^2 = \frac{\mu_h^2 + \mu_s^2}{2} \pm \frac{\mu_h^2 - \mu_s^2}{2} \sqrt{1 + y^2}$$

The Simplest Extension, Cont'd

Mass matrix



New topologies

$$\mu_h^2 \equiv \frac{\partial^2 V}{\partial h^2} = 2\bar{\lambda}_0 v_0^2$$

$$\mu_s^2 \equiv \frac{\partial^2 V}{\partial s^2} = b_3 x_0 + 2b_4 x_0^2 - \frac{a_1 v_0^2}{4x_0}$$

$$\mu_{hs}^2 \equiv \frac{\partial^2 V}{\partial h \partial s} = (a_1 + 2a_2 x_0) v_0$$

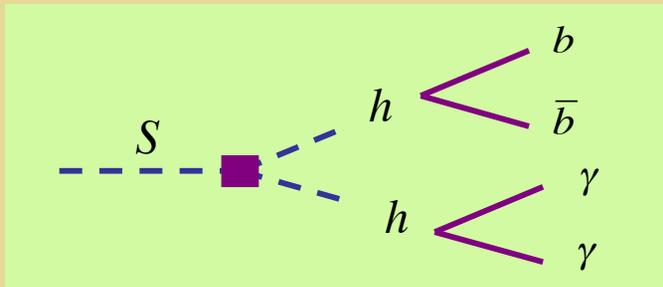
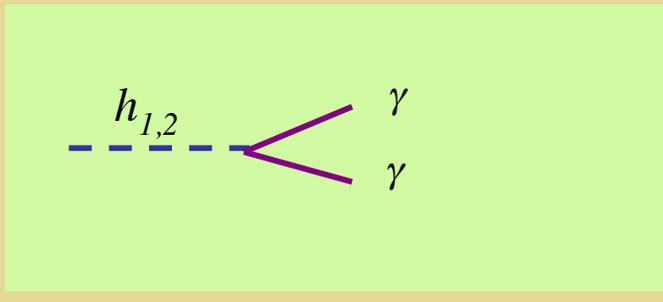
$$x_0 = \langle S^2 \rangle$$

$$\tan \theta = \frac{y}{1 + \sqrt{1 + y^2}}, \quad y \equiv \frac{\mu_{hs}^2}{\mu_h^2 - \mu_s^2}$$

$$m_{1,2}^2 = \frac{\mu_h^2 + \mu_s^2}{2} \pm \frac{\mu_h^2 - \mu_s^2}{2} \sqrt{1 + y^2}$$

The Simplest Extension, Cont'd

Mass matrix



$$\mu_h^2 \equiv \frac{\partial^2 V}{\partial h^2} = 2\bar{\lambda}_0 v_0^2$$

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$$x_0 = \langle S \rangle$$

$$\tan \theta = \frac{y}{1 + \sqrt{1 + y^2}}$$

$$y \equiv \frac{\mu_{hs}^2}{\mu_h^2 - \mu_s^2}$$

Stable S (dark matter)

- Tree-level Z_2 symmetry: $a_1=0$ to prevent s - h mixing and one-loop $s \rightarrow hh$
- $x_0=0$ to prevent h - s mixing & $s \rightarrow hh$

The Simplest Extension

DM Scenario

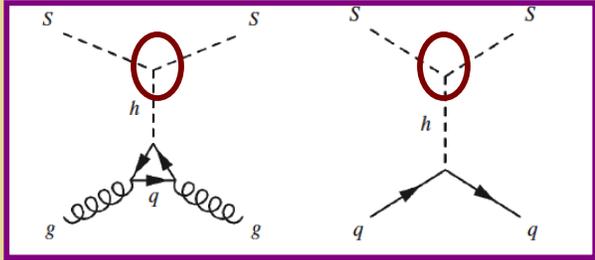
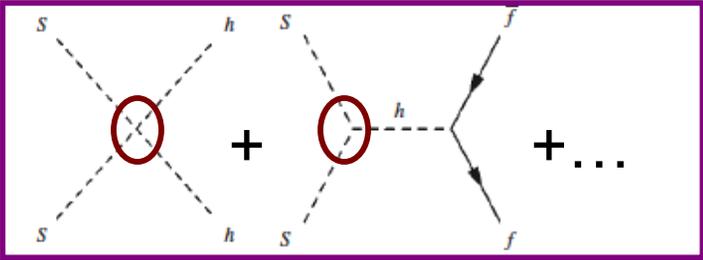
$$V_{\text{HS}} = \quad + \frac{a_2}{2} (H^\dagger H) S^2$$

The Simplest Extension

DM Scenario

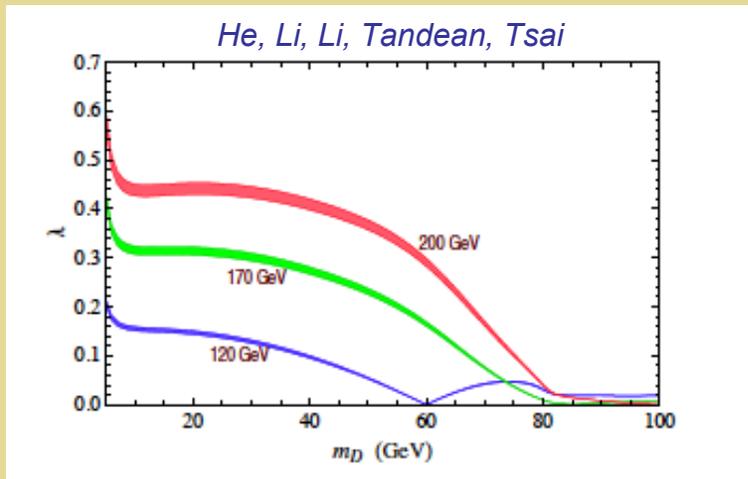
$$V_{\text{HS}} = \dots + \frac{a_2}{2} (H^\dagger H) S^2$$

$\Omega_{\text{DM}} \& \sigma_{\text{SI}}$

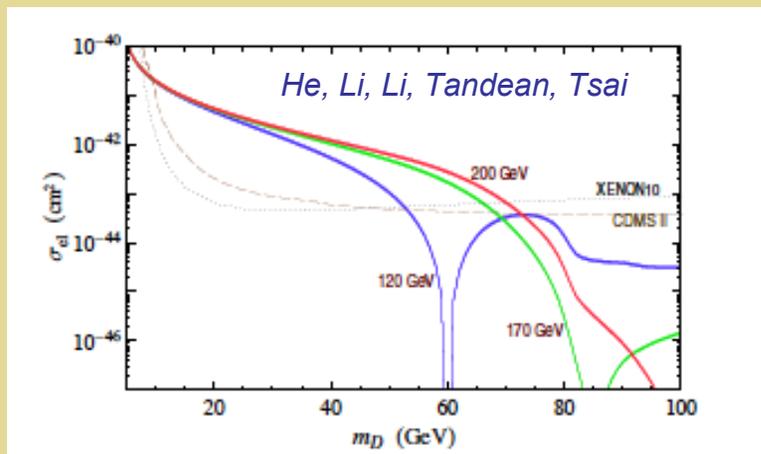


DM Phenomenology

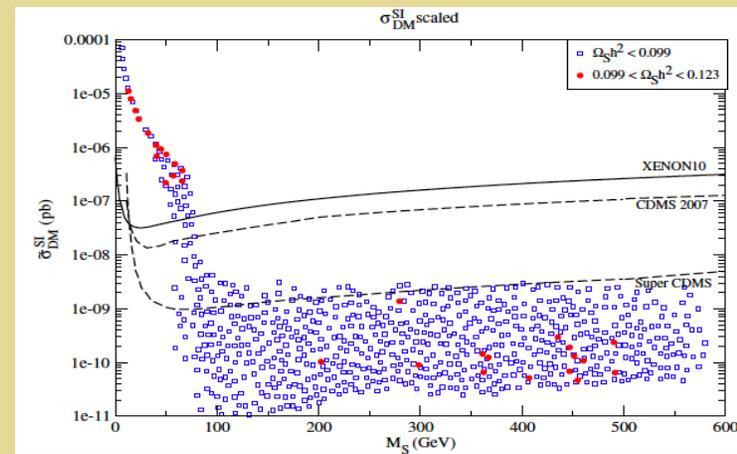
Relic Density



Direct Detection

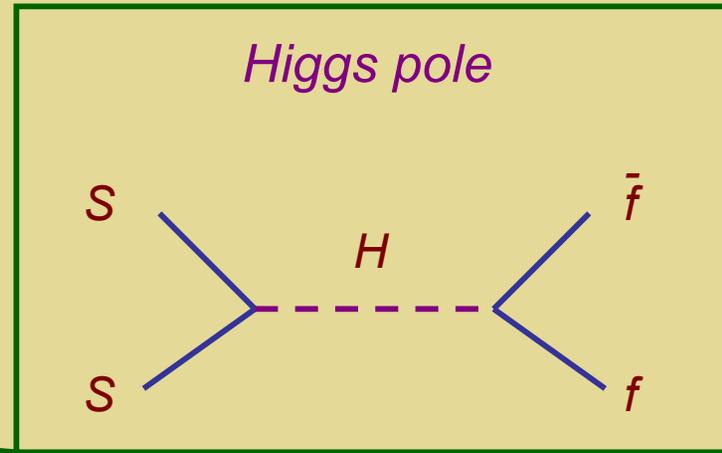
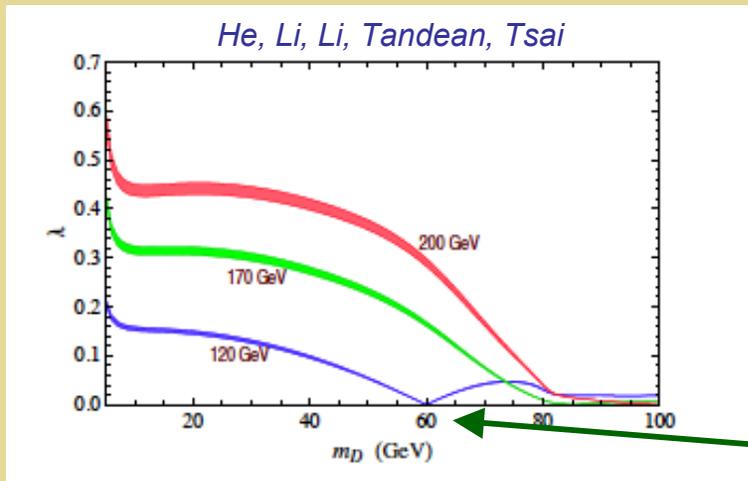


Barger, Langacker, McCaskey,
R-M, Shaughnessy

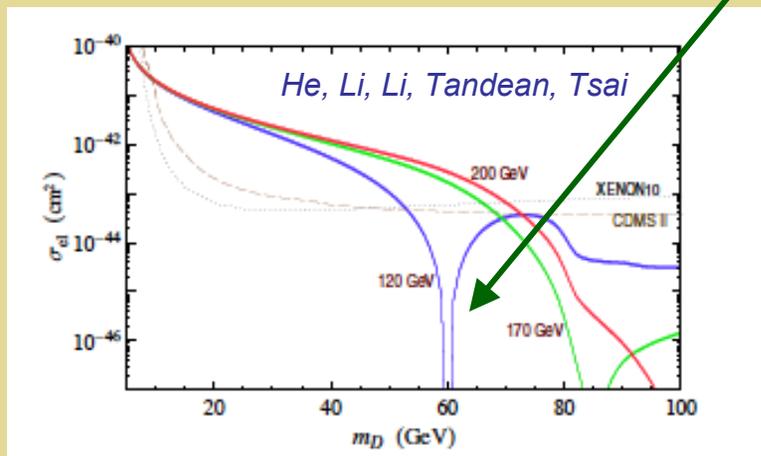


DM Phenomenology

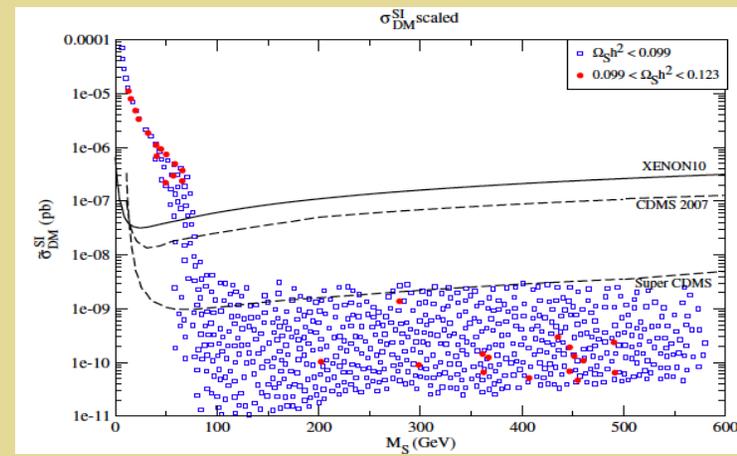
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Direct Detection

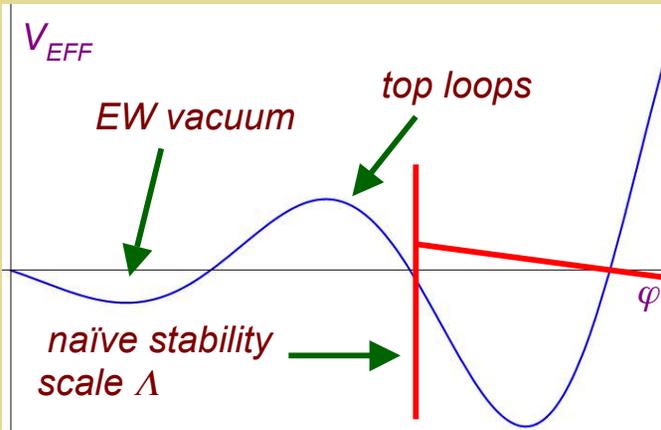


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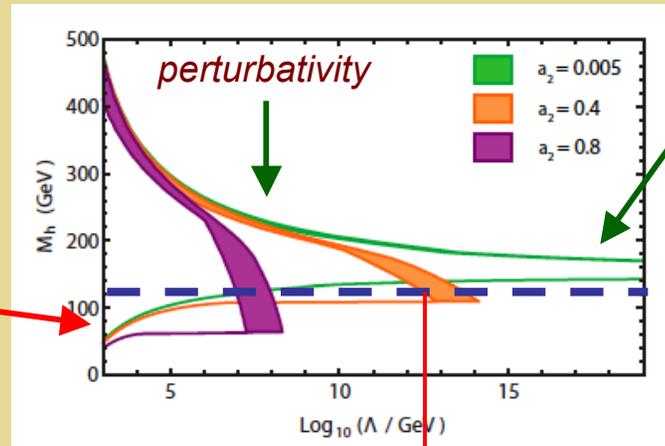


New Scalars EW Vacuum Stability

Preserving EW Min



"Funnel plot"



SM stability & pert'vity

SM + singlet: stable but non-pertur'tive

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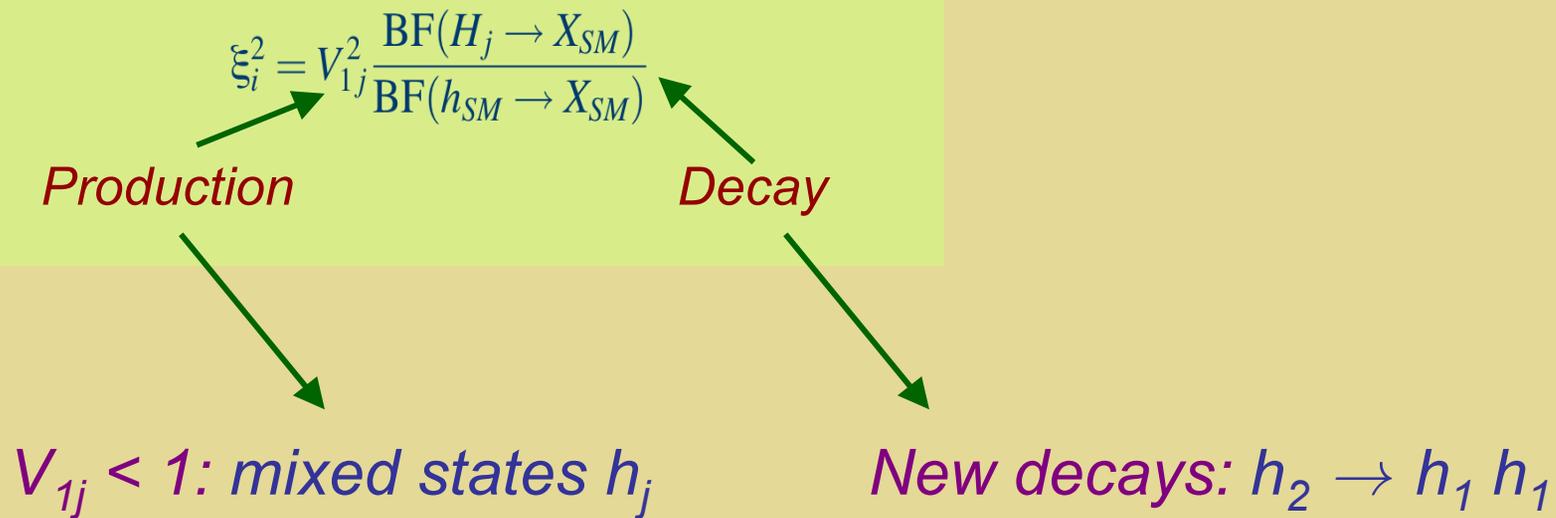
DM-H coupling

top loops

LHC & Higgs Phenomenology

LHC discovery potential

Signal Reduction Factor



Dark matter: no mixing \rightarrow states are h, S

New decays $h \rightarrow SS$ (invisible!) possible

LHC & Higgs Phenomenology

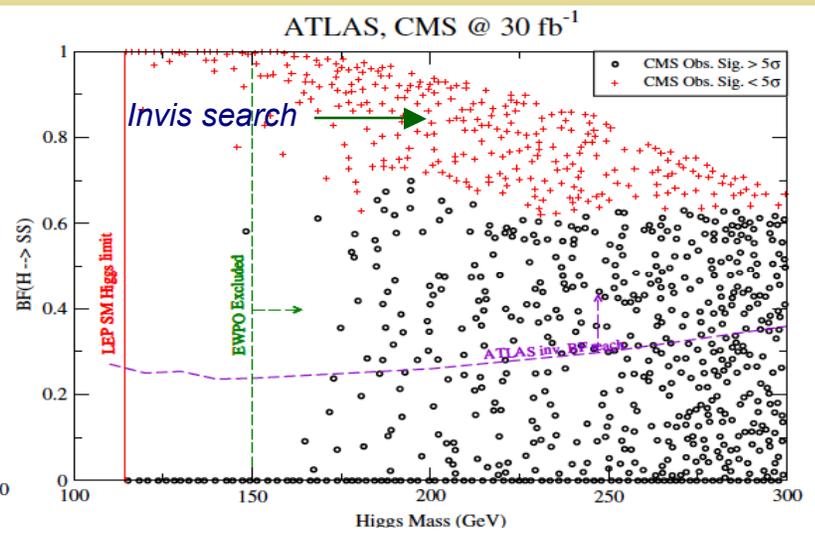
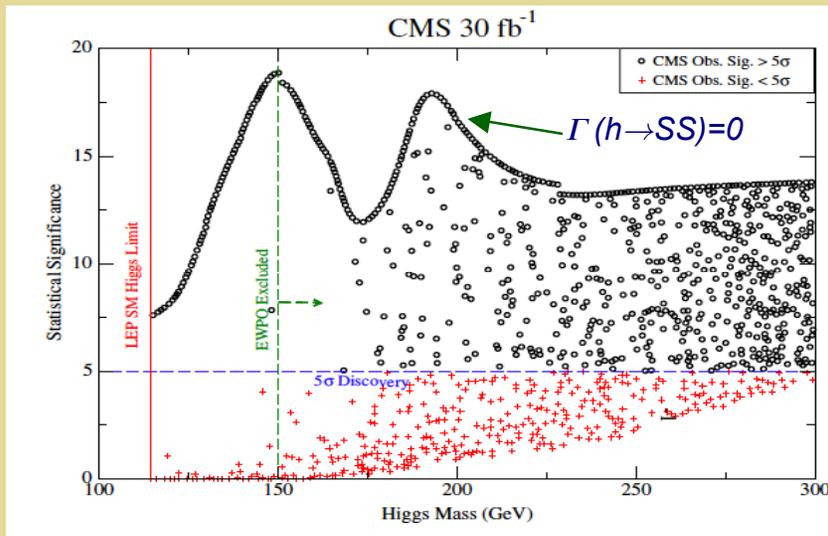
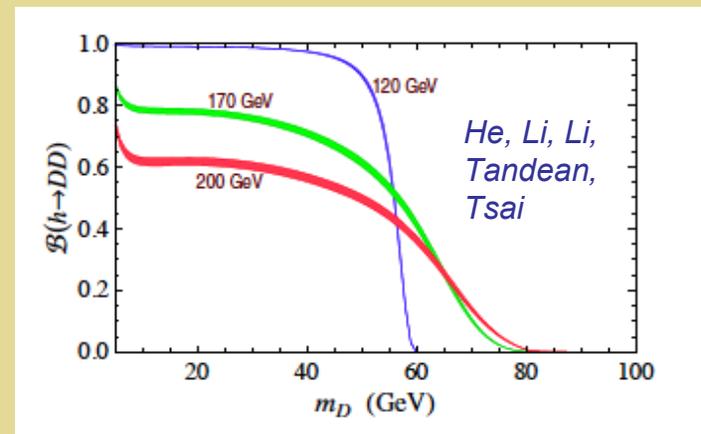
LHC discovery potential

Signal Reduction Factor

$$\xi_i^2 = V_{1j}^2 \frac{\text{BF}(H_j \rightarrow X_{SM})}{\text{BF}(h_{SM} \rightarrow X_{SM})}$$

Production Decay

Invisible decays



LHC & Higgs Phenomenology

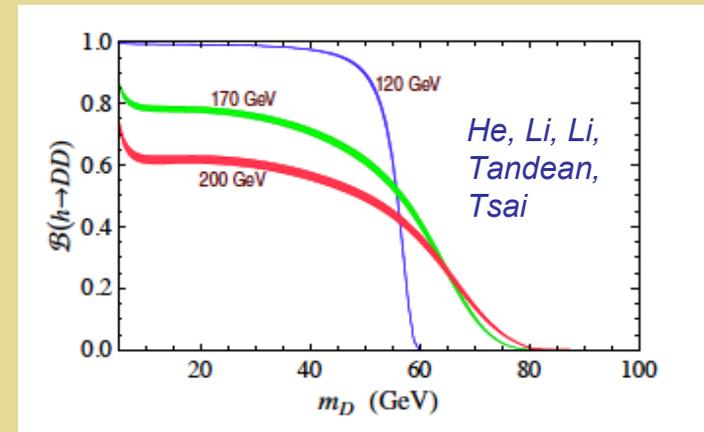
LHC discovery potential

Signal Reduction Factor

$$\xi_i^2 = V_{1j}^2 \frac{\text{BF}(H_j \rightarrow X_{SM})}{\text{BF}(h_{SM} \rightarrow X_{SM})}$$

Production Decay

Invisible decays



Dijet azimuthal distribution



Look for azimuthal shape change of primary jets (Eboli & Zeppenfeld '00)

LHC & Higgs Phenomenology

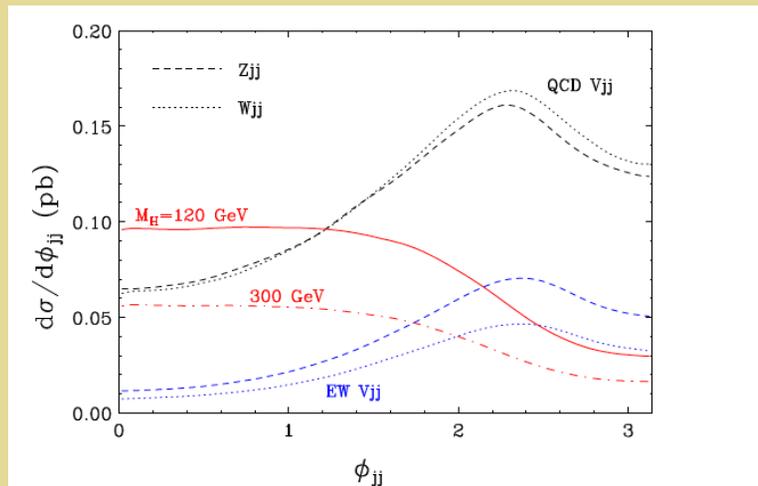
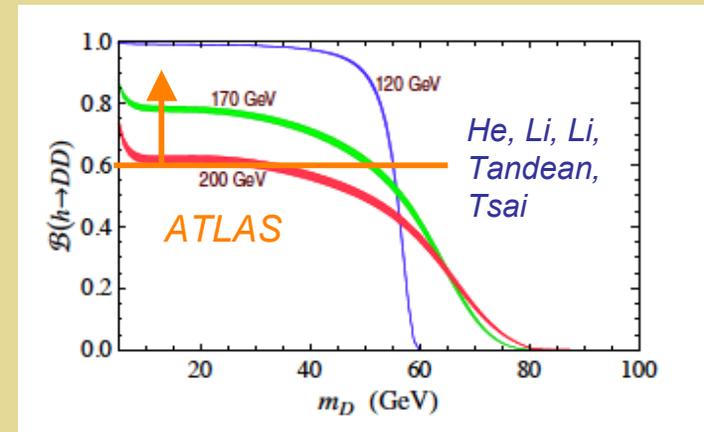
LHC discovery potential

Signal Reduction Factor

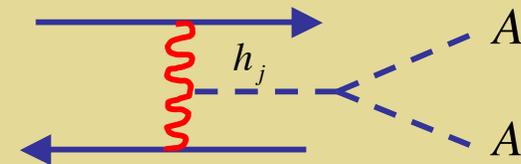
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Production Decay

Invisible decays

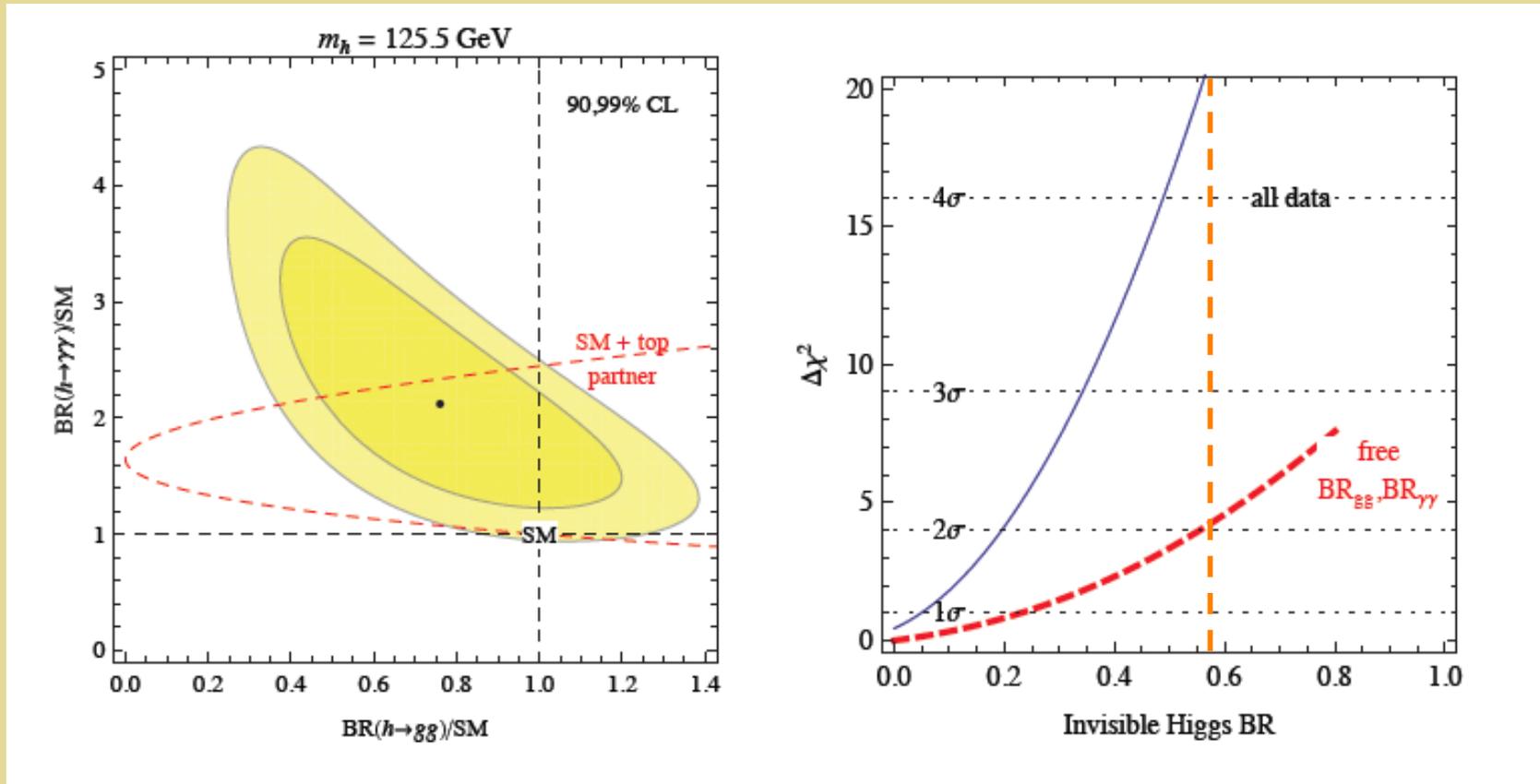


Dijet azimuthal distribution



Look for azimuthal shape change of primary jets (Eboli & Zeppenfeld '00)

LHC & Higgs Phenomenology



Giardino et al:
arXiv 1207:1347

Higgs Portal: Simple Scalar Extensions

<i>Extension</i>	<i>DOF</i>	<i>EWPT</i>	<i>DM</i>
<i>Real singlet</i>	<i>1</i>	<i>✓</i>	<i>✗</i>
<i>Real singlet</i>	<i>1</i>	<i>✗</i>	<i>✓</i>
<i>Complex Singlet</i>	<i>2</i>	<i>✓</i>	<i>✓</i>
<i>Real Triplet</i>	<i>3</i>	<i>✓</i>	<i>✓</i>

May be low-energy remnants of UV complete theory & illustrative of generic features

Complex Singlet: EWB & DM?

Barger, Langacker, McCaskey, R-M Shaugnessy

Spontaneously & softly broken global U(1) $\langle S \rangle \neq 0$

$$V_{HS} = \frac{\delta_2}{2} H^\dagger H |\tilde{S}|^2 = \frac{\delta_2}{2} H^\dagger H (S^2 + A^2)$$



Controls Ω_{CDM} , T_C , & H-S mixing

$$V_{\tilde{S}} = \frac{b_2}{2} |\tilde{S}|^2 + \frac{b_1}{2} \tilde{S}^2 + \text{c.c.} + \dots$$



Gives non-zero M_A

Complex Singlet: EWB & DM?

Barger, Langacker, McCaskey, R-M Shaugnessy

Consequences:

Three scalars: h_1, h_2 : mixtures of h & S

A : dark matter

Phenomenology:

- Produce h_1, h_2 w/ reduced σ

- Reduce BR ($h_j \rightarrow SM$)

- Observation of BR (invis)

- Possible obs of σ^{SI}

Complex Singlet: EWB & DM?

Barger, Langacker, McCaskey, R-M Shaugnessy

Consequences:

Impact EWPO

Three scalars:

h_1, h_2 : mixtures of h & S

A : dark matter

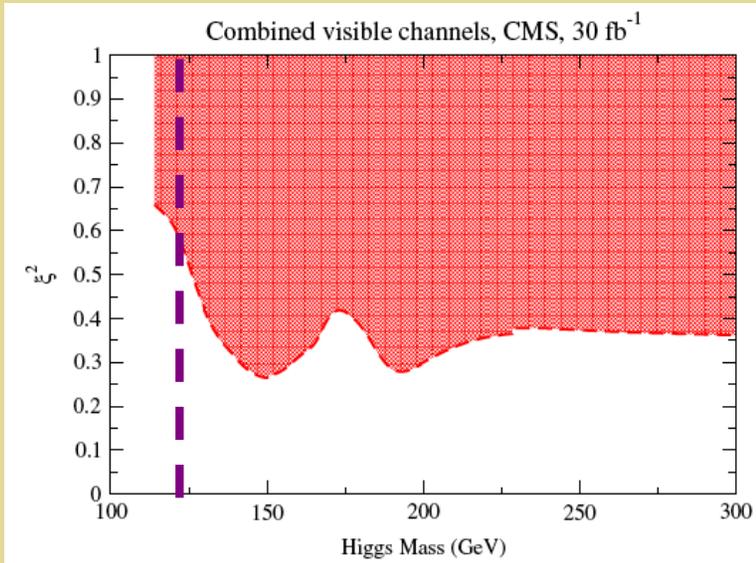
Phenomenology:

- *Produce h_1, h_2 w/ reduced σ*
- *Reduce BR ($h_j \rightarrow SM$)*
- *Observation of BR (invis)*
- *Possible obs of σ^{SI}*

Complex Singlet: LHC Discovery

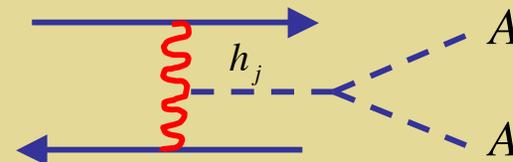
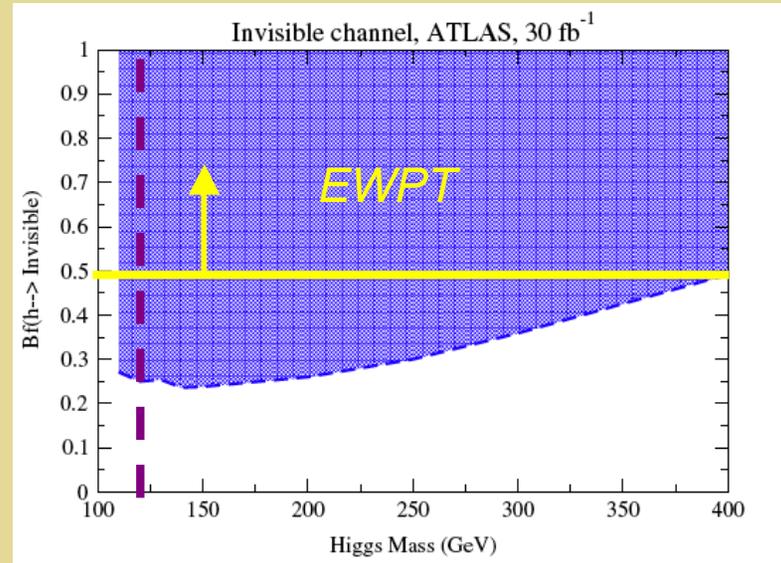
Barger, Langacker,
McCaskey, R-M, Shaughnessy

Traditional search: CMS

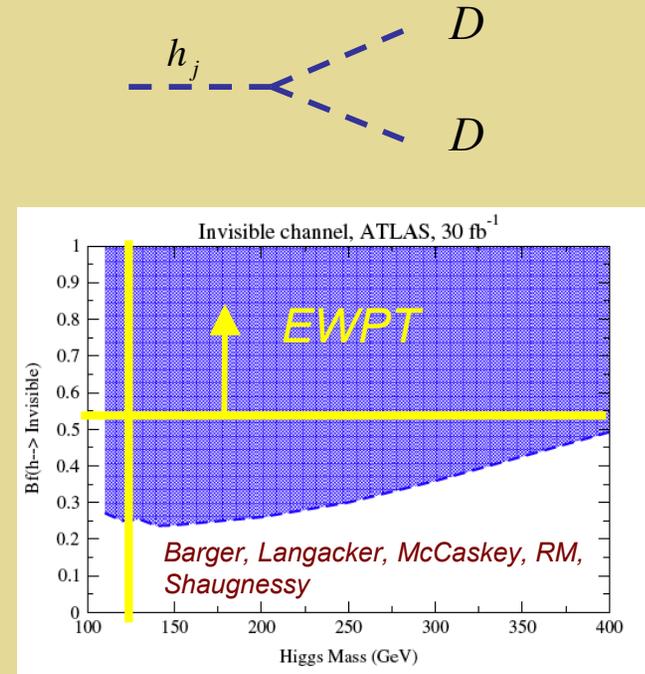
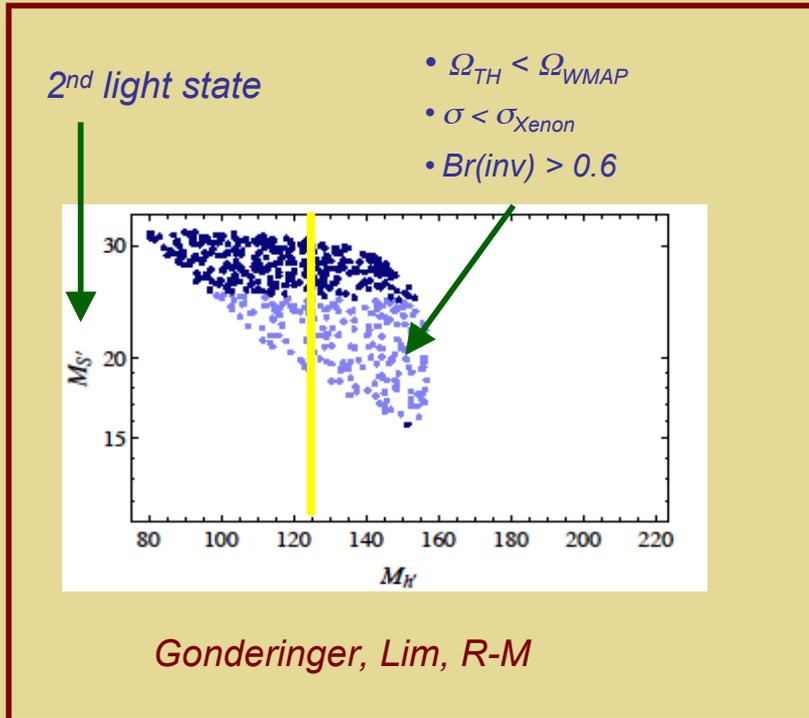


Single component case ($x_0 \neq 0$)

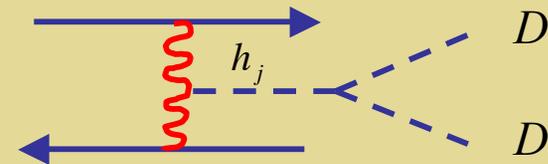
Invisible search: ATLAS



Complex Singlet: New Light State ?



Three scalars: $h_{1,2}$ (Higgs-like)
 D (dark matter)



LHC: WBF "Invisible decay" search

Higgs Portal: Simple Scalar Extensions

<i>Extension</i>	<i>DOF</i>	<i>EWPT</i>	<i>DM</i>
<i>Real singlet</i>	1	✓	✗
<i>Real singlet</i>	1	✗	✓
<i>Complex Singlet</i>	2	✓	✓
<i>Real Triplet</i>	3	✓	✓

Simplest non-trivial EW multiplet

Real Triplet

$\Sigma^0, \Sigma^+, \Sigma^-$

$\sim (1, 3, 0)$

Fileviez-Perez, Patel, Wang, R-M: PRD
79: 055024 (2009); 0811.3957 [hep-ph]

$$V_{H\Sigma} = \frac{a_1}{2} H^\dagger \Sigma H + \frac{a_2}{2} H^\dagger H \text{Tr} \Sigma^2$$

EWPT: $a_{1,2} \neq 0$ & $\langle \Sigma^0 \rangle \neq 0$

DM & EWPT: $a_1 = 0$ & $\langle \Sigma^0 \rangle = 0$

Real Triplet

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Small: ρ -param

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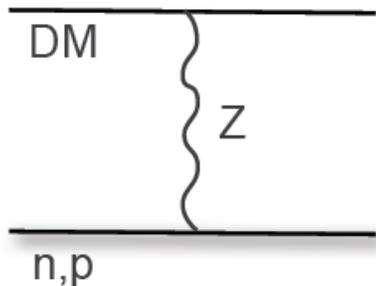
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R-M: 0811.3957 [hep-ph]

$$V_{H\Sigma} = \frac{a_2}{2} H^\dagger H \text{Tr} \Sigma^2$$

New feature: gauge interactions & direct detection



$$\sigma_n \sim \frac{g_1^2}{4\pi^2} \frac{m_{n,p}^2}{m_Z^4} \sim 10^{-39} \text{cm}^2$$

$$g_{\chi\chi Z} \propto 2 I_3 - 4 Q \sin^2 \theta_W$$

$$Y = 2 (Q - I_3)$$

→ *Want $Y = 0$*

Real Triplet

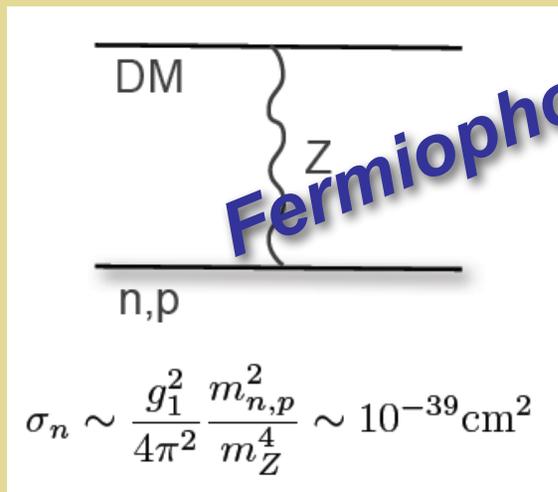
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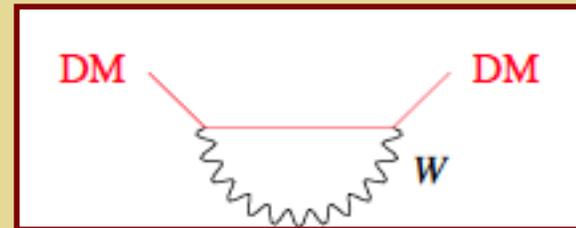
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Fileviez-Perez, Patel, Wang,
R-M: 0811.3957 [hep-ph]

$$V_{H\Sigma} = \frac{a_2}{2} H^\dagger H \text{Tr} \Sigma^2$$

EWPO

$$T_\Sigma \approx -\frac{1}{6\pi\hat{s}^2} \frac{(\Delta M)^2}{M_W^2}$$

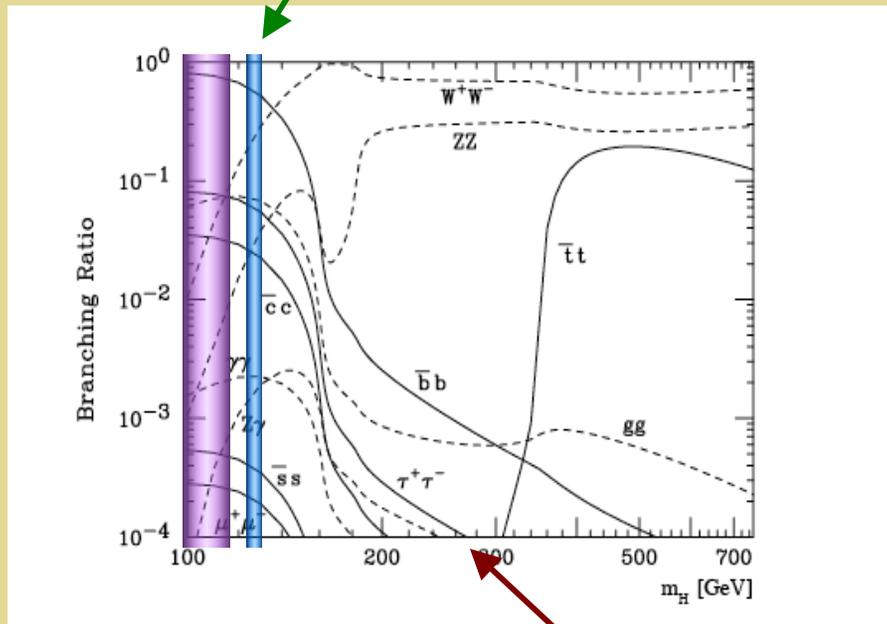


$$\alpha S \sim \Pi_{3\gamma} = 0$$

$$\Delta M = 166 \text{ MeV}$$

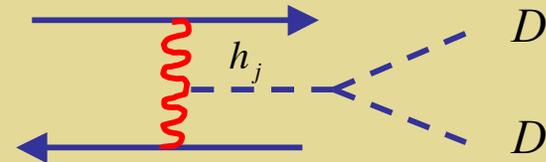
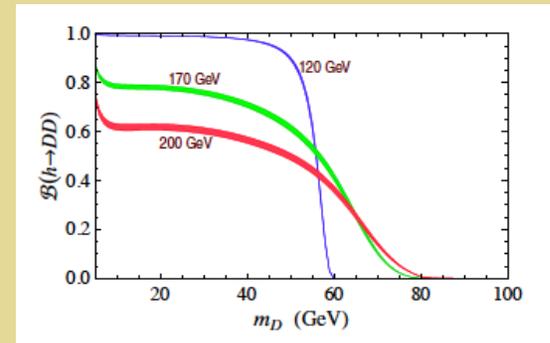
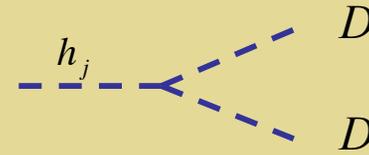
LHC Phenomenology

SM Branching Ratios ?



Four scalars: h_1 (Higgs-like)
 Σ^0 (dark matter)
 Σ^+, Σ^- (new states)

Fileviez-Perez, Patel, R-M, Wang



LEP: $M_\Sigma > 100 \text{ GeV} \rightarrow$
 $BR(\text{invis}) = 0$

Production

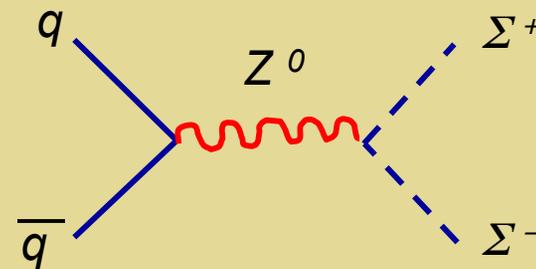
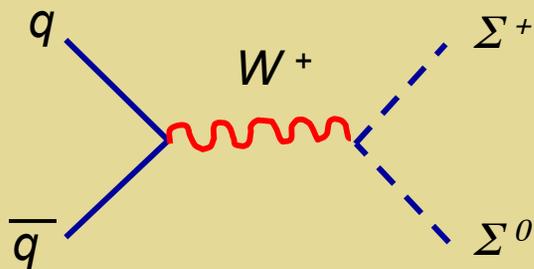
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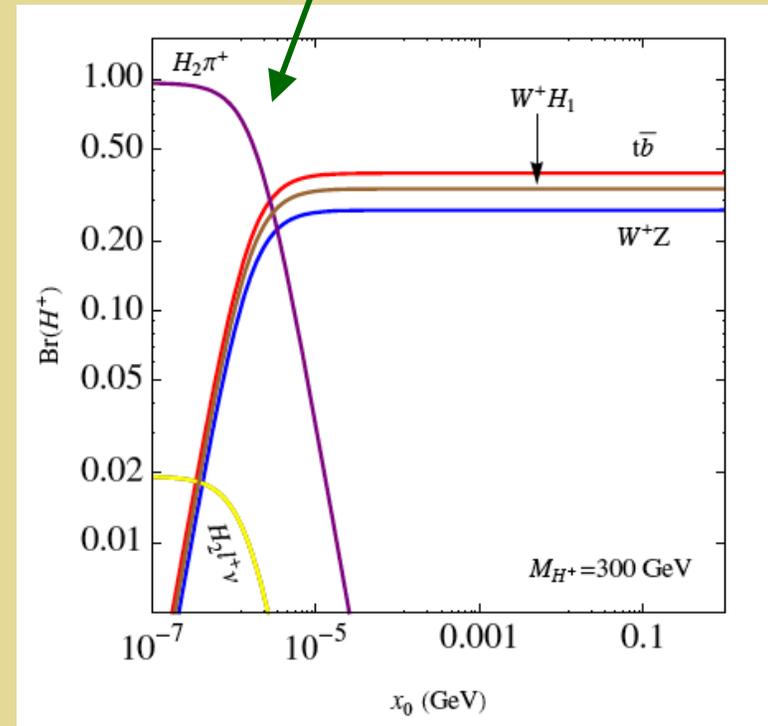
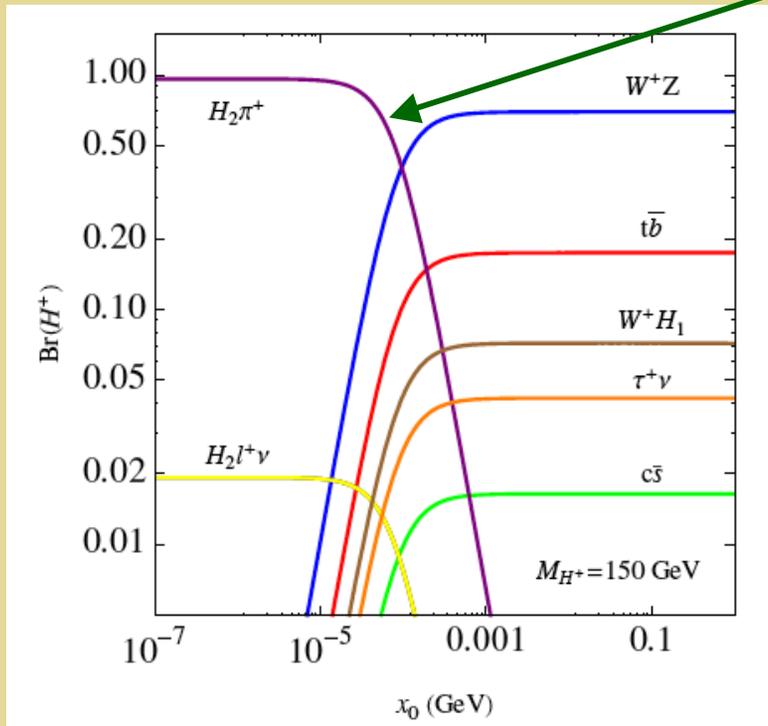


EW cross sections w/ charged states + $\cancel{\tau}$

Decays : Charged BRs

Charged: x_0 dependence

$H^\pm \rightarrow H_2 W^{\pm*} \rightarrow H_2 \pi^\pm$ Pure gauge



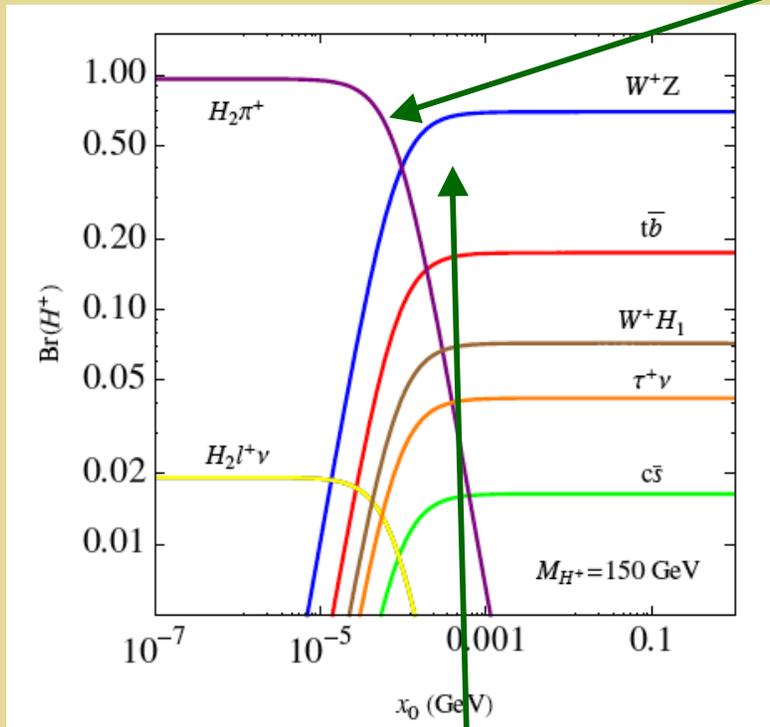
**Below WZ
Threshold**

DM: $x_0 = 0$

**Above WZ
Threshold**

Decays : Charged BRs

Charged: x_0 dependence

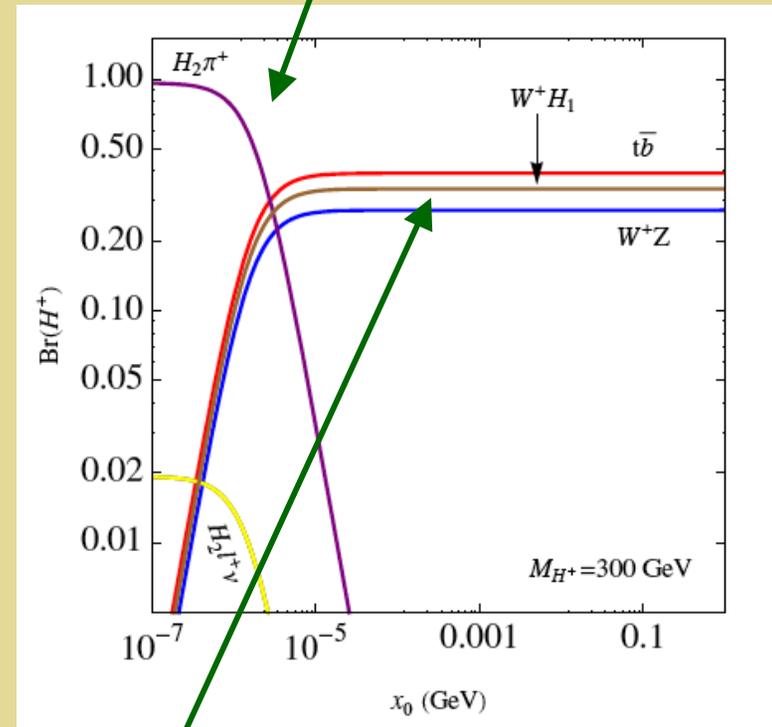


**Below WZ
Threshold**

$H^\pm \rightarrow W^\pm Z, f_1 \bar{f}_2$

x_0 dependent

$H^\pm \rightarrow H_2 W^{\pm*} \rightarrow H_2 \pi^\pm$ Pure gauge



**Above WZ
Threshold**

Real Triplet : DM Search

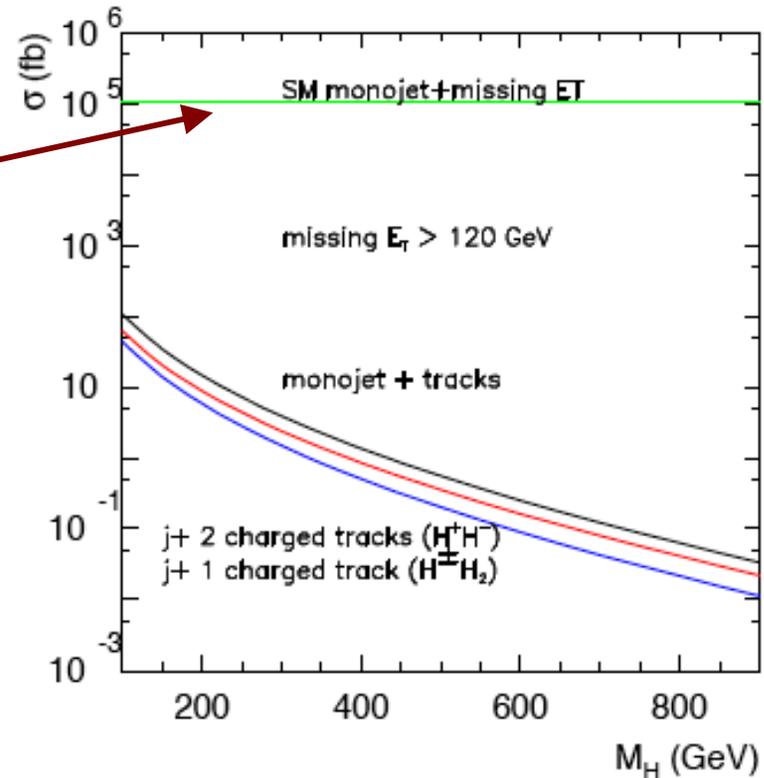
Basic signature: Charged track disappearing after ~ 5 cm

$$x_0 = 0 : H^\pm \rightarrow H_2 \pi^\pm$$

$$q\bar{q} \rightarrow W^{\pm*} \rightarrow H^\pm H_2 \quad q\bar{q} \rightarrow Z^*, \gamma^* \rightarrow H^+ H^-$$

Trigger: Monojet (ISR) + large \cancel{E}_T

SM Background:
QCD jZ and jW w/
 $Z \rightarrow \nu\nu$ & $W \rightarrow l\nu$



Real Triplet : DM Search

Basic signature: Charged track disappearing after ~ 5 cm

$$x_0 = 0 : H^\pm \rightarrow H_2 \pi^\pm$$

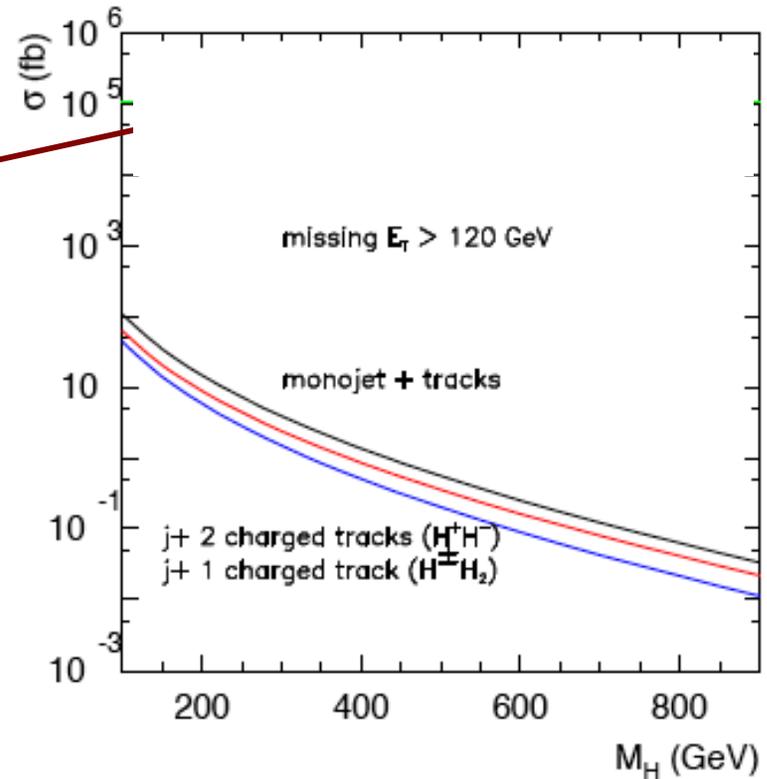
after ~ 5 cm

$$q\bar{q} \rightarrow W^{\pm*} \rightarrow H^\pm H_2 \quad q\bar{q} \rightarrow Z^*, \gamma^* \rightarrow H^+ H^-$$

Trigger: Monojet
(ISR) + large \cancel{E}_T

SM Background:
QCD jZ and jW w/
 $Z \rightarrow \nu\nu$ & $W \rightarrow l\nu$

Cuts: large \cancel{E}_T
hard jet
One 5cm track



Real Triplet : DM Search

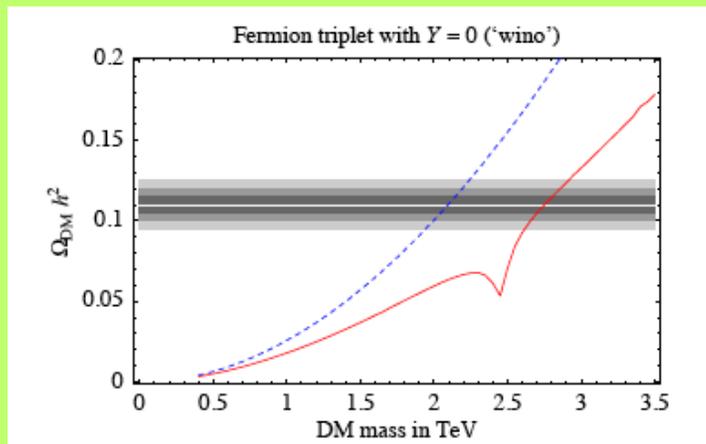
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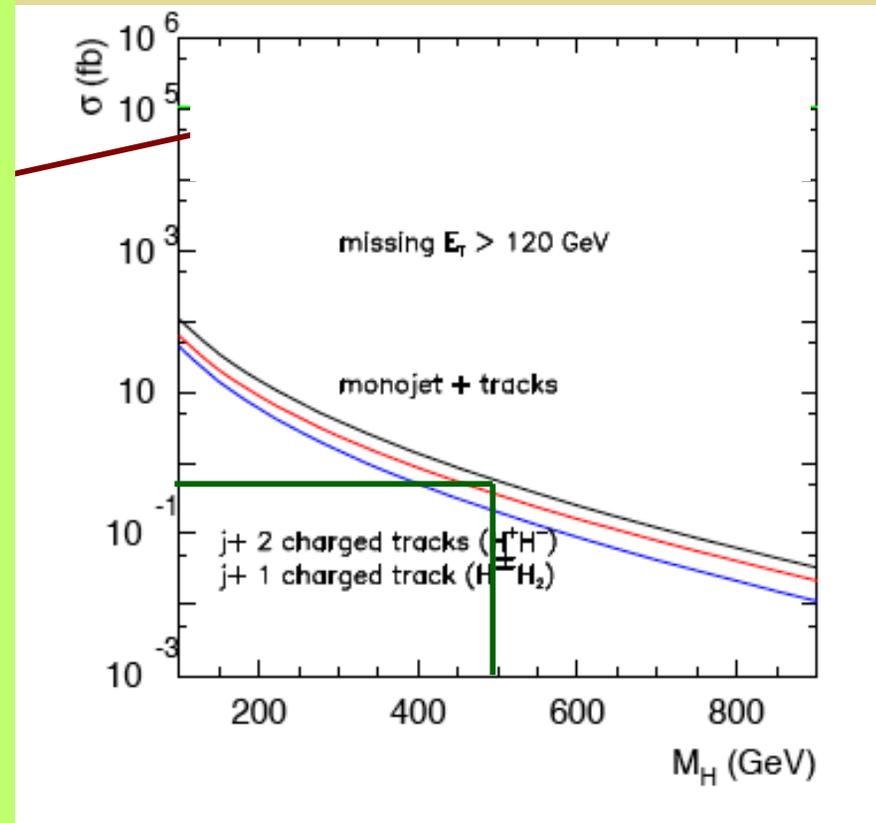
$$q\bar{q} \rightarrow W^{\pm*} \rightarrow H^\pm H_2 \quad q\bar{q} \rightarrow Z^*, \gamma^* \rightarrow H^+ H^-$$

Cirelli et al:



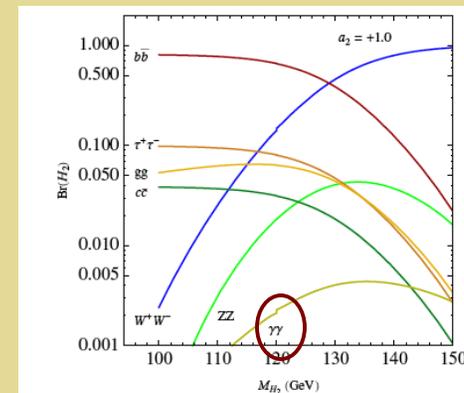
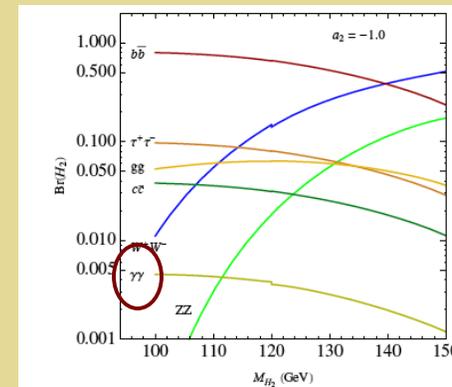
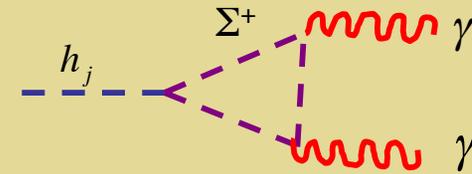
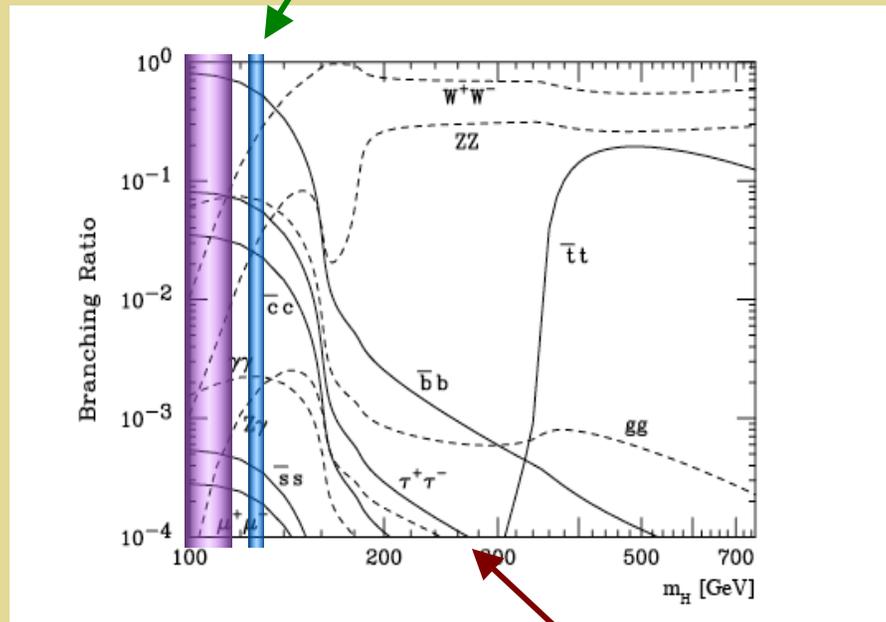
$$M_\Sigma = 500 \text{ GeV:}$$

$$\Omega_\Sigma / \Omega_{\text{CDM}} \sim 0.1$$



Higgs Diphoton Decay

SM Branching Ratios ?

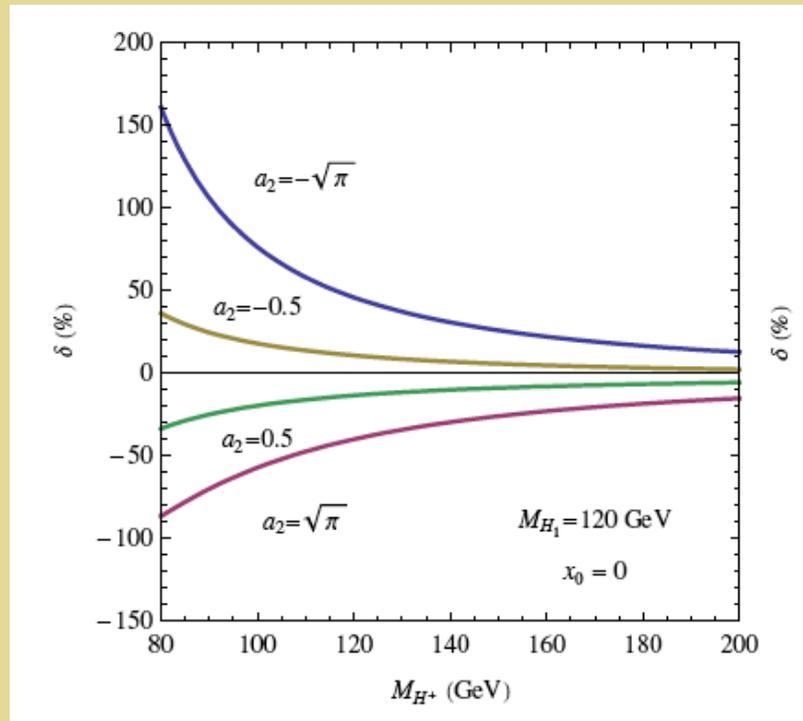


Four scalars: h_1 (Higgs-like)
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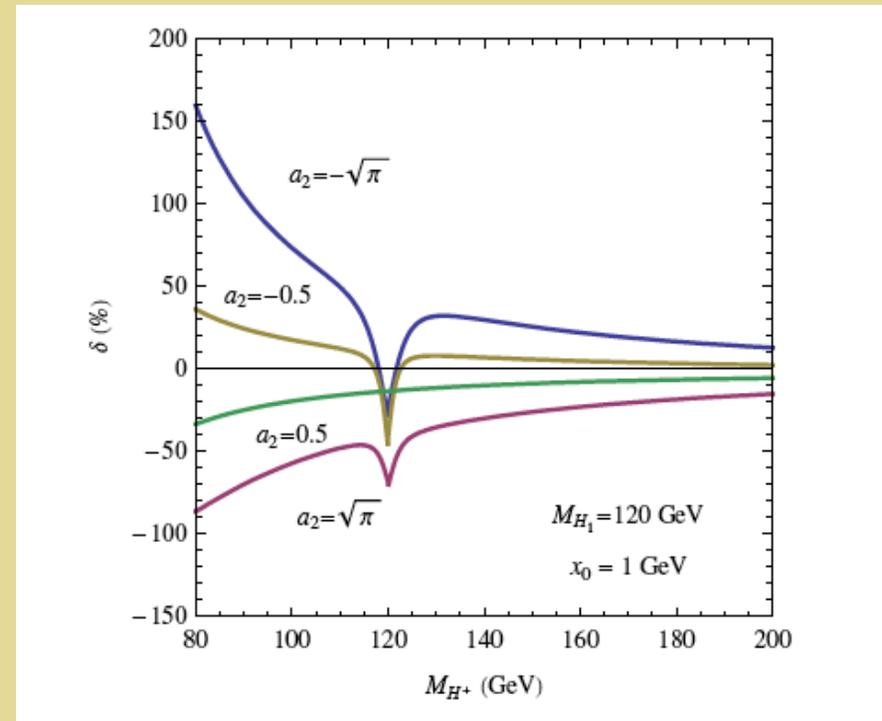
Fileviez-Perez, Patel, R-M, Wang

Real Triplet : H_1 Diphoton Decays

Neutral SM-like Higgs: H^+ loops and BR ($H_1 \rightarrow \gamma\gamma$)



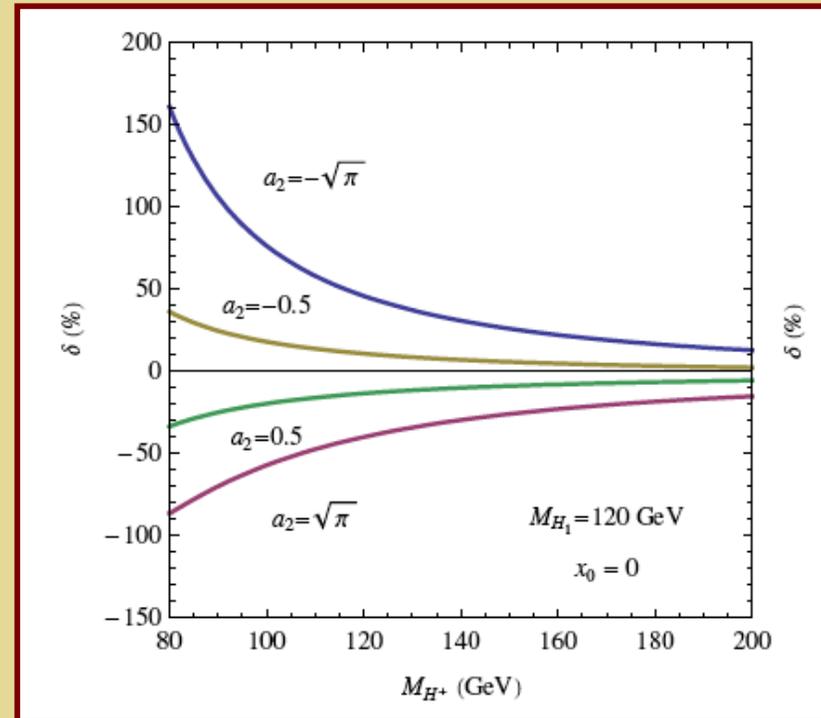
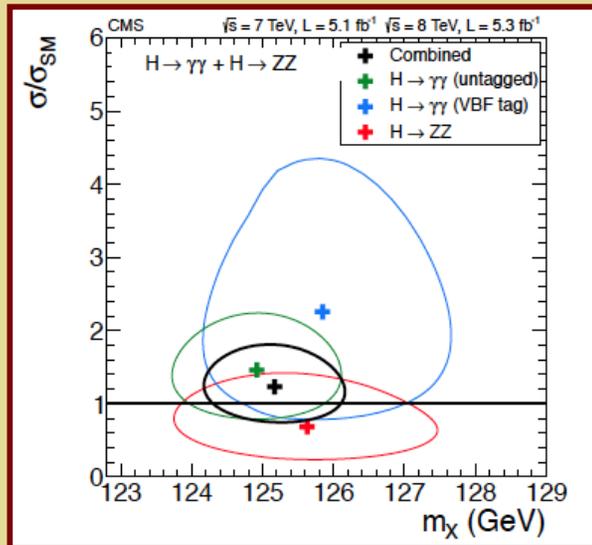
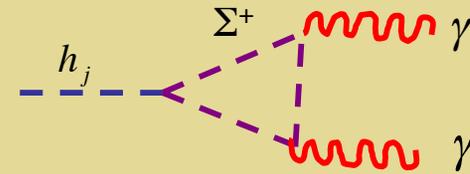
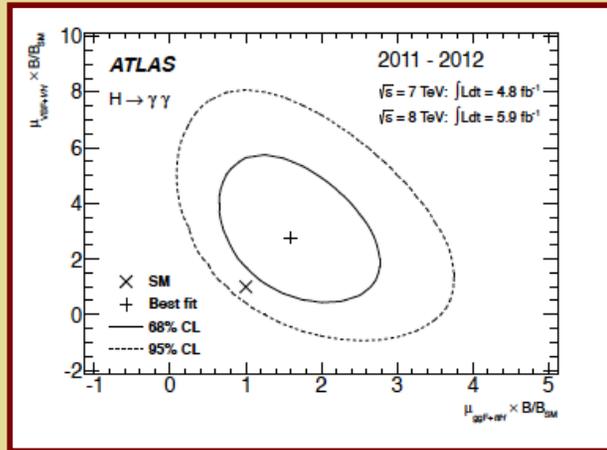
$x_0 = 0$: DM &
EWPT case



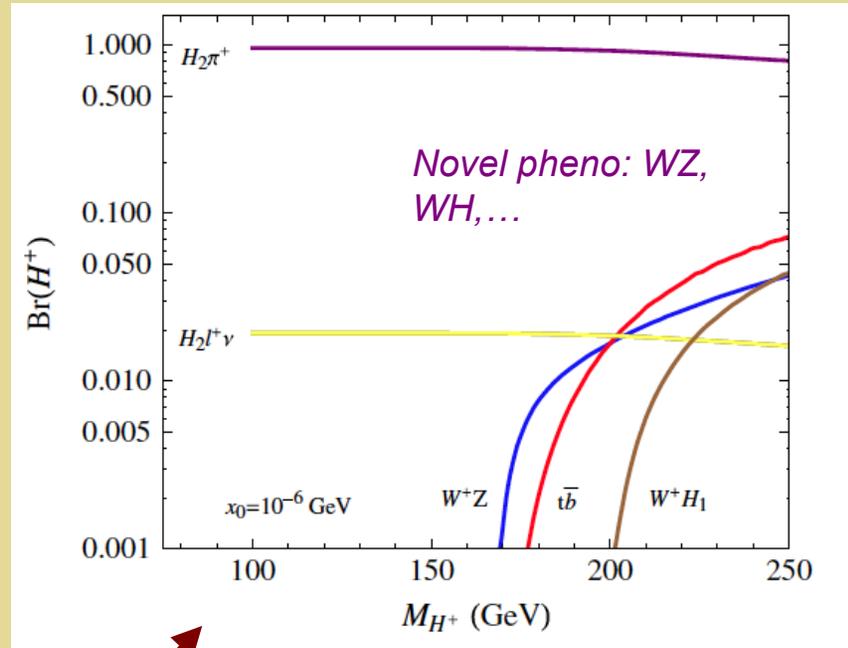
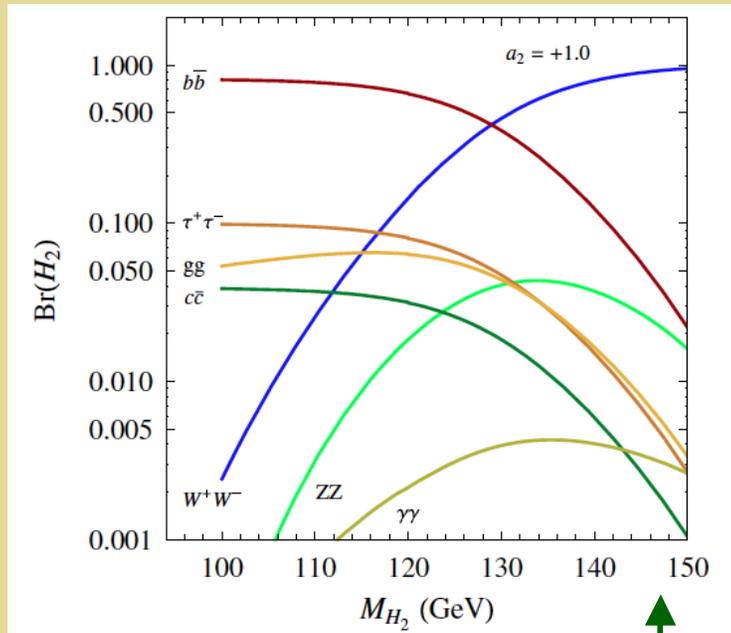
$x_0 > 0$: EWPT case

Higgs Diphoton Decays

LHC: $H \rightarrow \gamma\gamma$



EWPT: Heavier States ?



Four scalars: h_1 (Higgs-like)

$h_2 \sim \Sigma^0$ (unstable)

Σ^+, Σ^- (new states)

Topologies:

$bb \tau \nu$

$\gamma\gamma \tau \nu \dots$

VI. Discussion

Higher Dim Multiplets

Cirelli, Forengeno, Strumia '05

Quantum numbers			DM can	DM mass	$m_{\text{DM}^\pm} - m_{\text{DM}}$	Events at LHC	σ_{SI} in
$\text{SU}(2)_L$	$\text{U}(1)_Y$	Spin	decay into	in TeV	in MeV	$\int \mathcal{L} dt = 100/\text{fb}$	10^{-45} cm^2
2	1/2	0	EL	0.54 ± 0.01	350	$320 \div 510$	0.2
2	1/2	1/2	EH	1.1 ± 0.03	341	$160 \div 330$	0.2
3	0	0	HH^*	2.0 ± 0.05	166	$0.2 \div 1.0$	1.3
3	0	1/2	LH	2.4 ± 0.06	166	$0.8 \div 4.0$	1.3
3	1	0	HH, LL	1.6 ± 0.04	540	$3.0 \div 10$	1.7
3	1	1/2	LH	1.8 ± 0.05	525	$27 \div 90$	1.7
4	1/2	0	HHH^*	2.4 ± 0.06	353	$0.10 \div 0.6$	1.6
4	1/2	1/2	(LHH^*)	2.4 ± 0.06	347	$5.3 \div 25$	1.6
4	3/2	0	HHH	2.9 ± 0.07	729	$0.01 \div 0.10$	7.5
4	3/2	1/2	(LHH)	2.6 ± 0.07	712	$1.7 \div 9.5$	7.5
5	0	0	(HHH^*H^*)	5.0 ± 0.1	166	$\ll 1$	12
5	0	1/2	—	4.4 ± 0.1	166	$\ll 1$	12
7	0	0	—	8.5 ± 0.2	166	$\ll 1$	46

See also Cai, He, R-M, Tsai '11;
Kumericki et al '12:

Neutrino portal w/ $\phi \sim (1, 6, -1/2)$



Rich collider
pheno w/ σ_{EW}

Is it a Scalar ?

<i>Extension</i>	<i>DOF</i>	<i>EWPT</i>	<i>DM</i>
<i>Real singlet</i>	<i>1</i>	<i>✓</i>	<i>✗</i>
<i>Real singlet</i>	<i>1</i>	<i>✗</i>	<i>✓</i>
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<i>Real Triplet</i>	<i>3</i>	<i>✓</i>	<i>✓</i>

Is it a Scalar ?

<i>Extension</i>	<i>DOF</i>	<i>EWPT</i>	<i>DM</i>
<i>Real singlet</i>	<i>1</i>	<i>✓</i>	<i>✗</i>
<i>Real singlet</i>	<i>1</i>	<i>✗</i>	<i>✓</i>
<i>Complex Singlet</i>	<i>2</i>	<i>✓</i>	<i>✓</i>
<i>Real Triplet</i>	<i>3</i>	<i>✓</i>	<i>✓</i>



*Mixed Higgs-like states and/or modified BRs:
signal reduction ξ , invisible search...*

Is it a Scalar ?

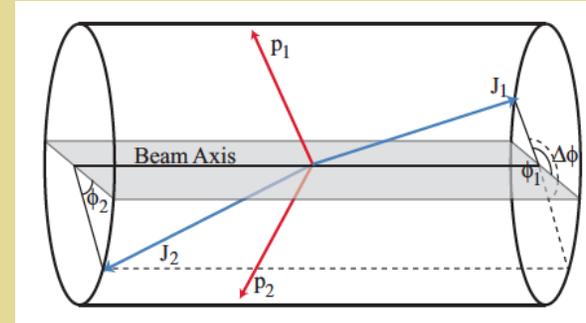
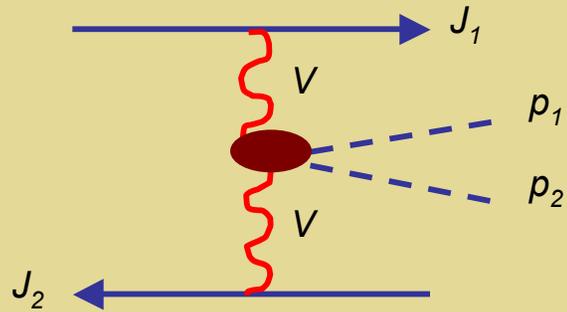
<i>Extension</i>	<i>DOF</i>	<i>EWPT</i>	<i>DM</i>
<i>Real singlet</i>	1	✓	✗
<i>Real singlet</i>	1	✗	✓
<i>Complex Singlet</i>	2	✓	✓
<i>Real Triplet</i>	3	✓	✓



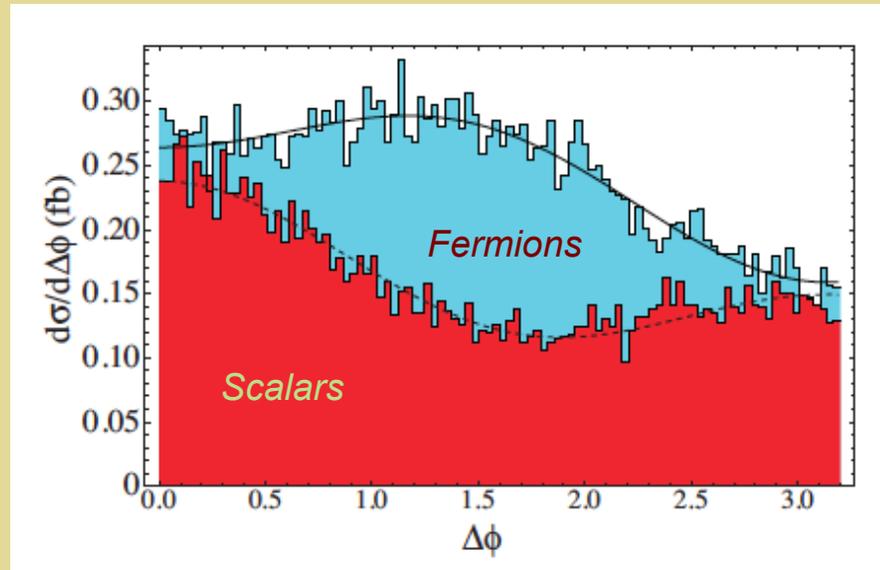
Could be fermionic triplet [e.g., Buckley, Randall, Shuve, JHEP 1105 (2011) 97]. How to distinguish?

Di-Jet Correlations

Buckley, R-M JHEP
1109 (2011) 094



*R-hadrons
from gg fusion*



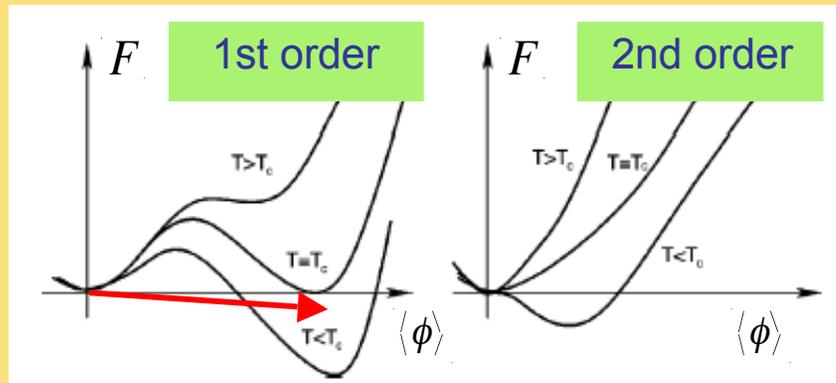
Summary

- *Observation of ~ 125 GeV Higgs-like scalar makes Higgs portal particularly interesting window on a number of questions: EW vacuum stability, gauge hierarchy, dark matter, EW phase transition...*
- *Higgs portal presents a rich array of possible signatures: modified Higgs properties, new (heavy) states, new event topologies associated w/ mixing and/or EW cross sections*

Back Up Slides

Electroweak Phase Transition

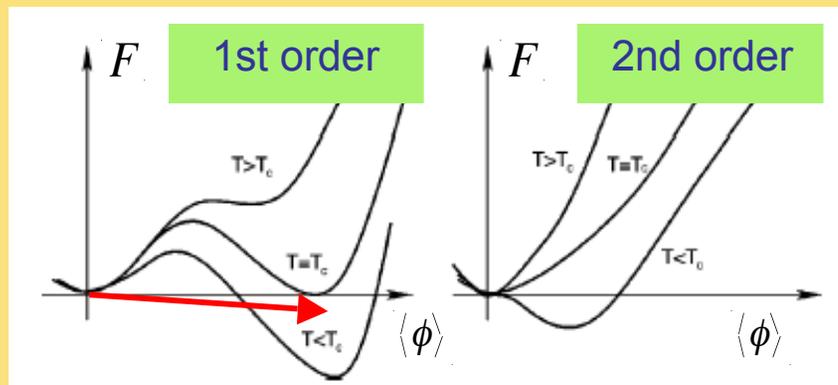
EW Phase Transition: New Scalars



Increasing m_h \longrightarrow

\longleftarrow New scalars

EW Phase Transition: New Scalars

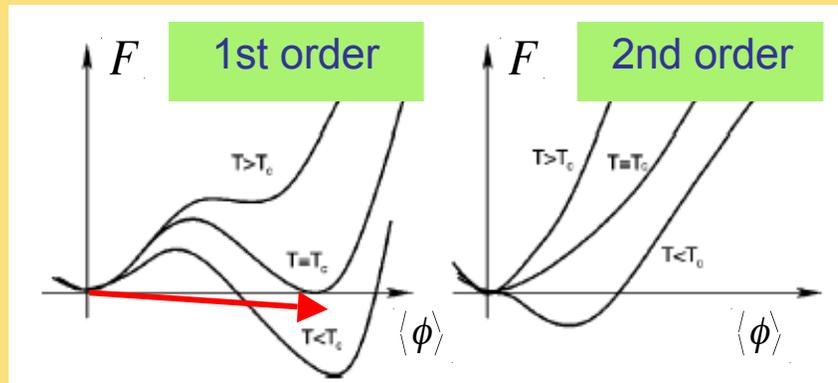


Increasing m_h \longrightarrow

\longleftarrow *New scalars*

Baryogenesis
Gravity Waves
Scalar DM
LHC Searches

EW Phase Transition: New Scalars



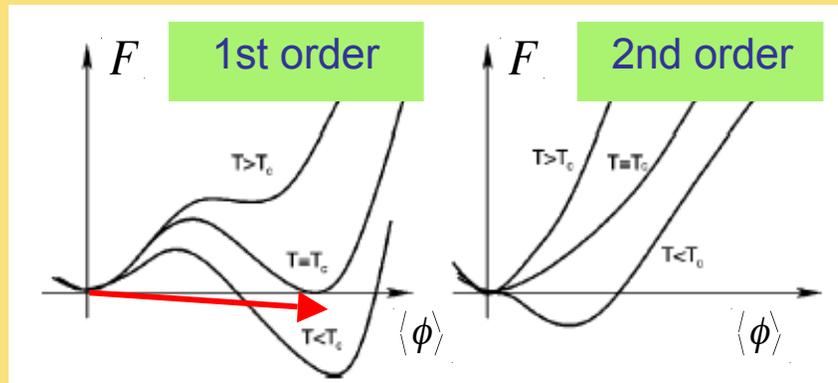
"Strong" 1st order EWPT

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EW Phase Transition: New Scalars



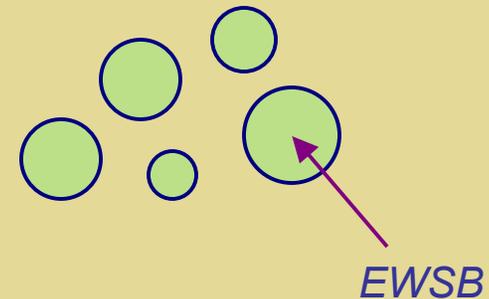
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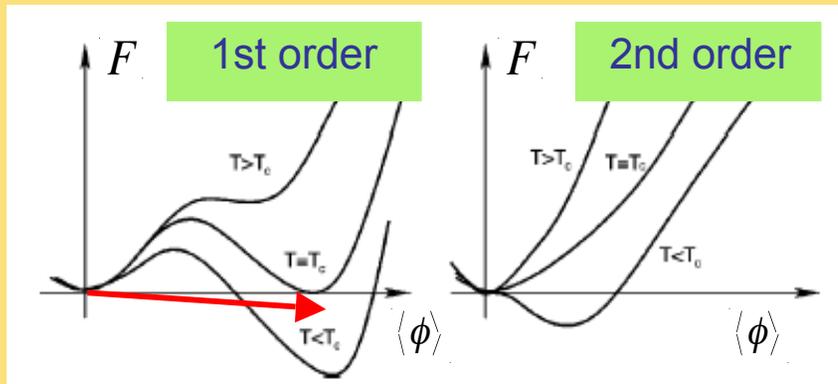
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“Strong” 1st order EWPT

Bubble nucleation



EW Phase Transition: New Scalars



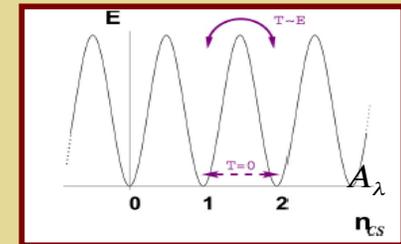
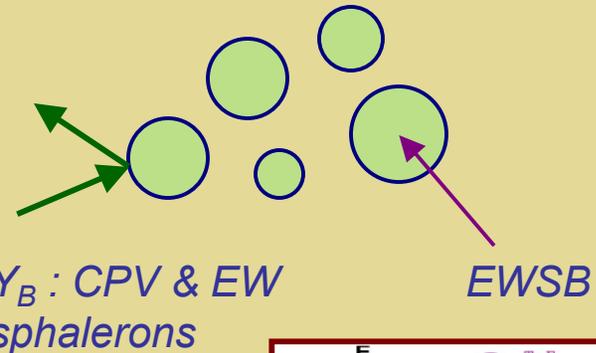
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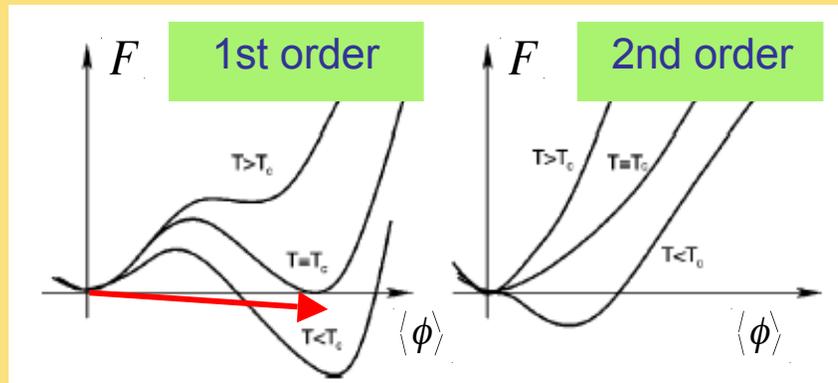
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EW Phase Transition: New Scalars



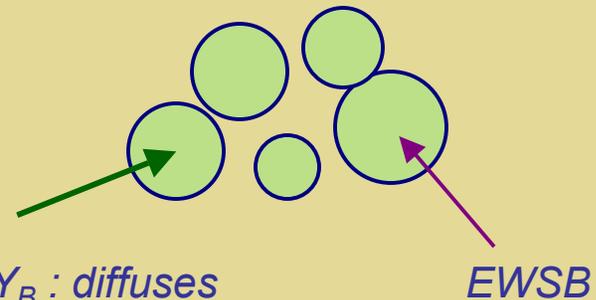
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Baryogenesis
Gravity Waves
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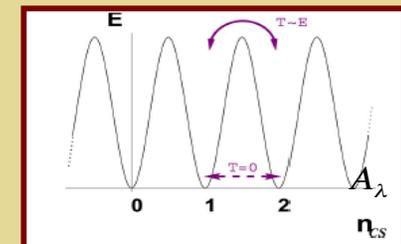
“Strong” 1st order EWPT

Bubble nucleation

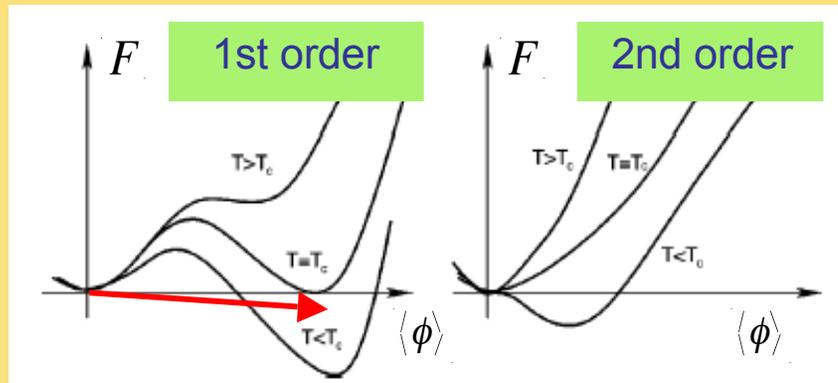


Y_B : diffuses into interiors

EWSB



EW Phase Transition: New Scalars



Increasing m_h \longrightarrow

\longleftarrow New scalars

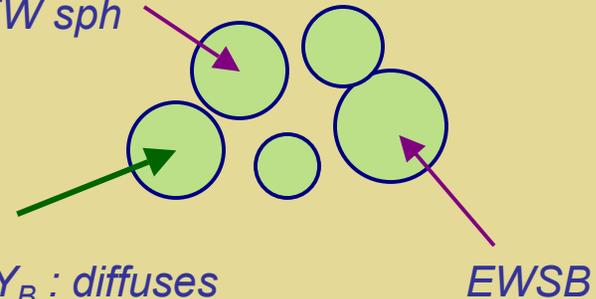
Baryogenesis
Gravity Waves
Scalar DM
LHC Searches

“Strong” 1st order EWPT

Preserve
 $Y_B^{initial}$

Bubble
nucleation

Quench
EW sph

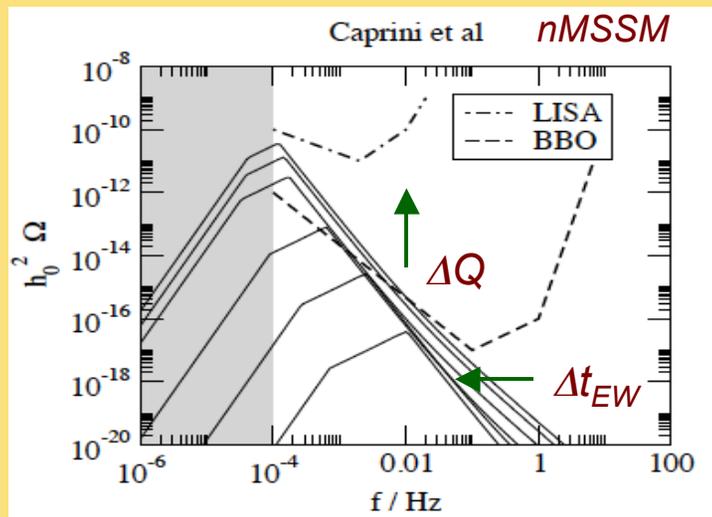
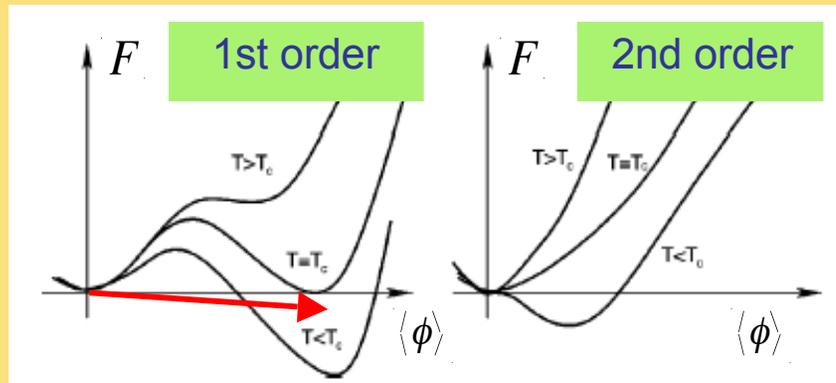


Y_B : diffuses
into interiors

EWSB

$T_C, E_{sph}, S_{tunnel} \longleftrightarrow F(\phi)$

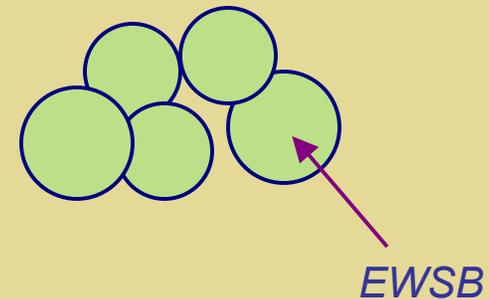
EW Phase Transition: New Scalars



"Strong" **1st order EWPT**

Detonation & turbulence

Bubble nucleation



GW Spectra:
 ΔQ & Δt_{EW}

$\longleftrightarrow F(\phi)$